

A complex visualization of the cosmic web, showing a dense network of filaments and nodes. The filaments are colored in vibrant shades of green, yellow, and orange, while the nodes are marked with black dots. The background is a dark, textured field of green and blue, suggesting the distribution of dark matter and gas in the universe.

the Cosmic Web:

Lecture 2b: Lagrangian Dynamics, Multistream Structure & Caustic Skeleton

Rien van de Weijgaert
Excsm Cosmology Summerschool, Haapsalu, July 2025

Cosmic Web

**Dynamics & Formation:
Program**

Cosmic Web – Formation & Dynamics

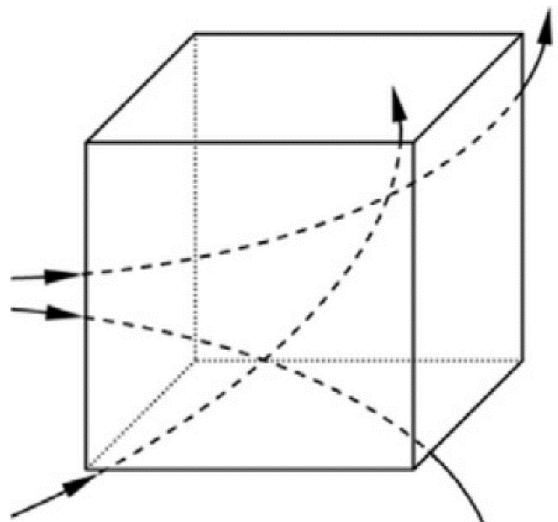
- **Forces & Strains** - **Observational Manifestations**
- **the Mechanism** - **Gravitational Instability**
- **Anisotropic Collapse** - **Formation of filaments and walls**
- **Weaving the Web** - **Connection Clusters, Filaments and Walls**
- **Dynamical Inventory** - **Forces & Tides in the Cosmic Web**
- **Phase Space Dynamics** - **Phase Space & Multistream structure**
- **Lagrangian Dynamics** - **Zeldovich formalism**
- **Hierarchical Formation** - **from small to the Megaparsec Cosmic Web**
- **Anisotropy & Hierarchy** - **the Adhesion formalism**
- **Caustic Skeleton** - **analytical formalism cosmic web**

Lagrangian View

Structure Formation

Eulerian - Lagrangian

Eulerian

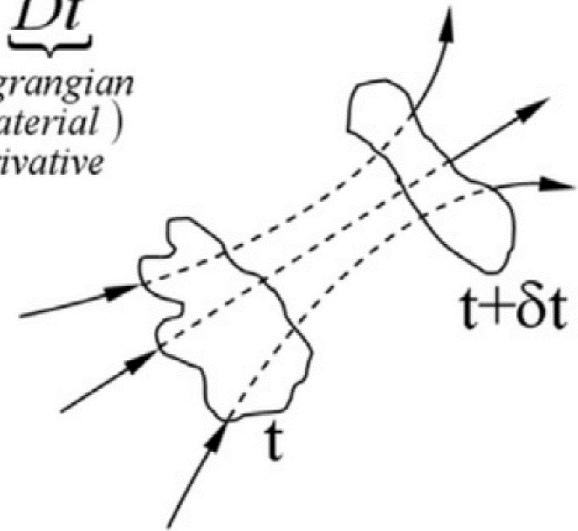


Spatially fixed
volume element

$$\underbrace{\frac{\partial}{\partial t} + \mathbf{v} \cdot \nabla}_{\text{Eulerian derivative}} =$$

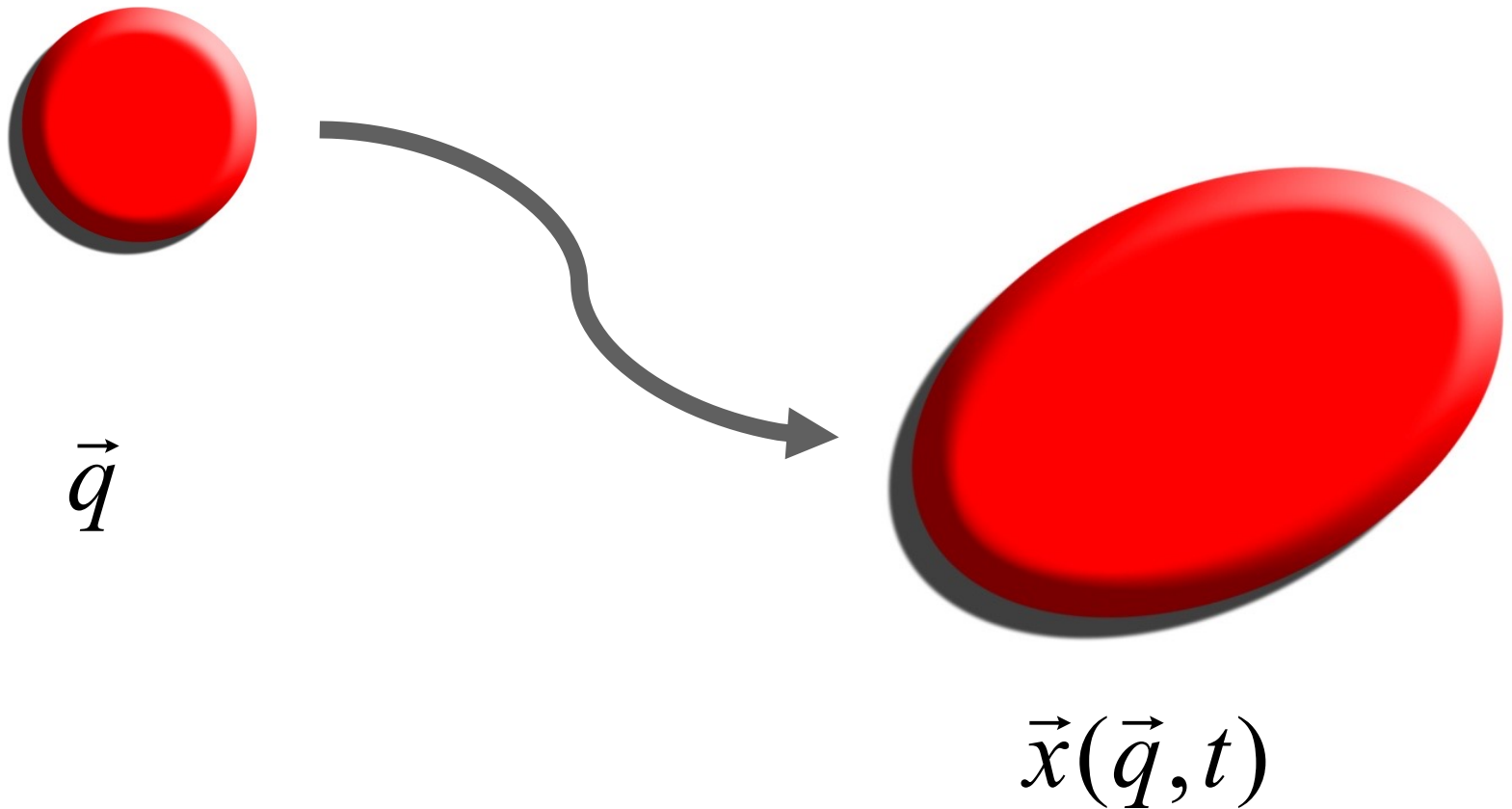
$$\underbrace{\frac{D}{Dt}}_{\text{Lagrangian (Material) derivative}}$$

Lagrangian

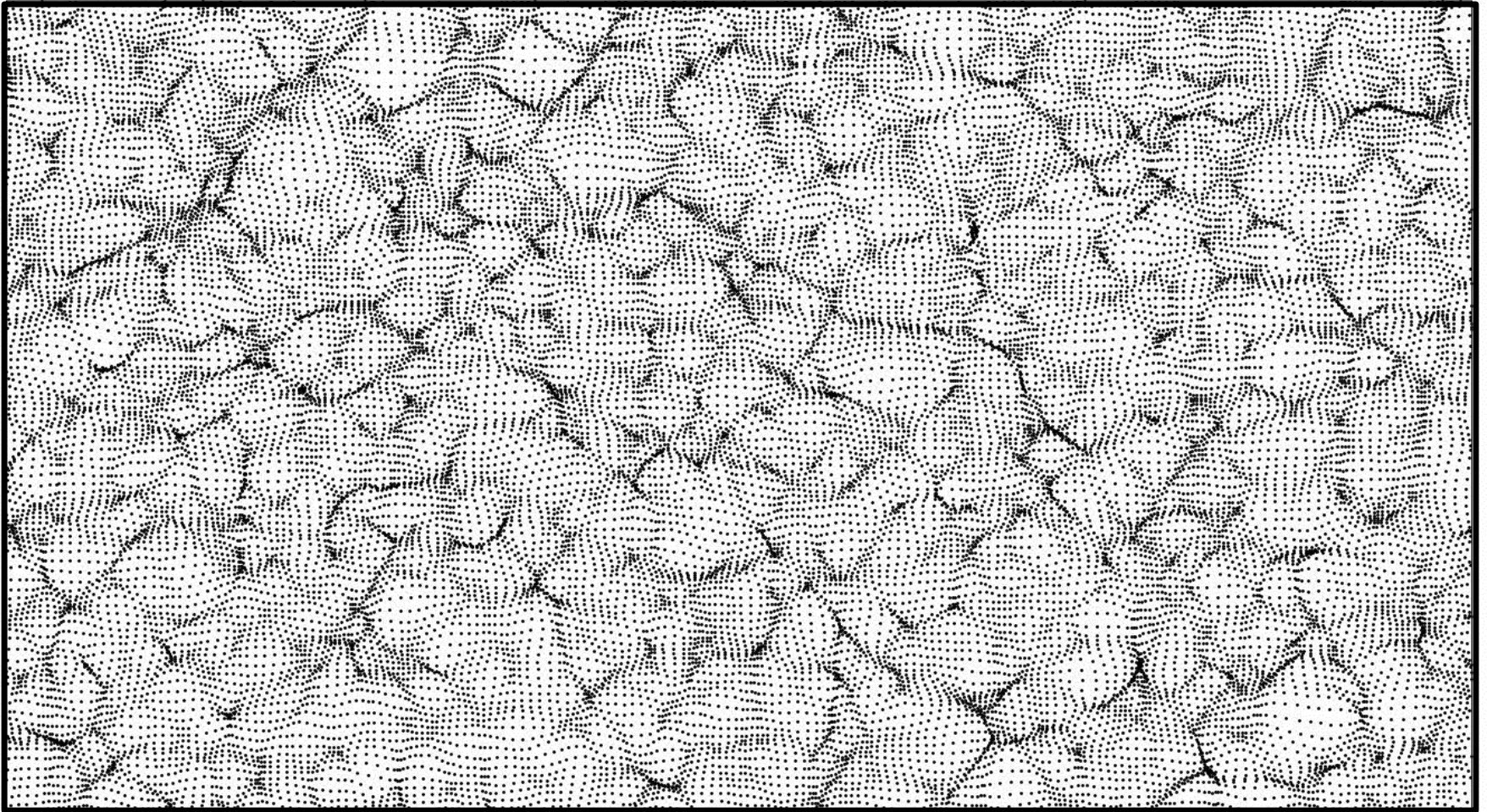


Following the motion
of the fluid element

Lagrangian - Eulerian



Simulation – Discrete Particles



Simulation – Mass Elements



Cosmic Web

**Phase-Space Dynamics
&
Multistream Regions**

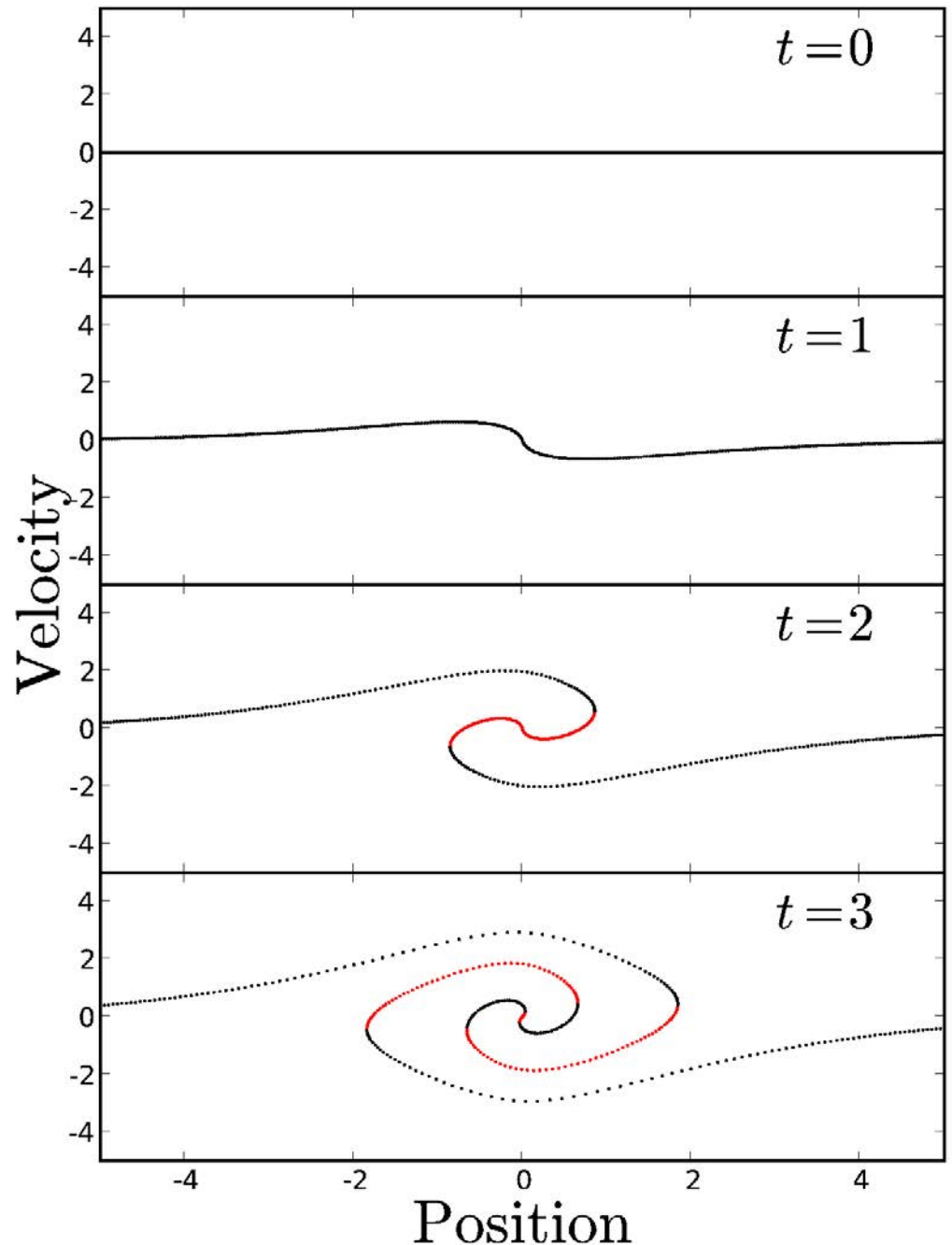
Phase Space Evolution

Dark Matter Phase Space sheet:

3-D structure projection of a
folding DM phase space sheet
In 6-D phase space

- Shandarin 2010, 2011
- Neyrinck et al. 2011, 2012
- Origami
- Abel et al. 2011

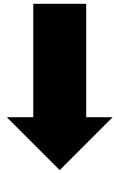
Evolving matter distribution in
position-velocity space – 1D



Phase Space Evolution

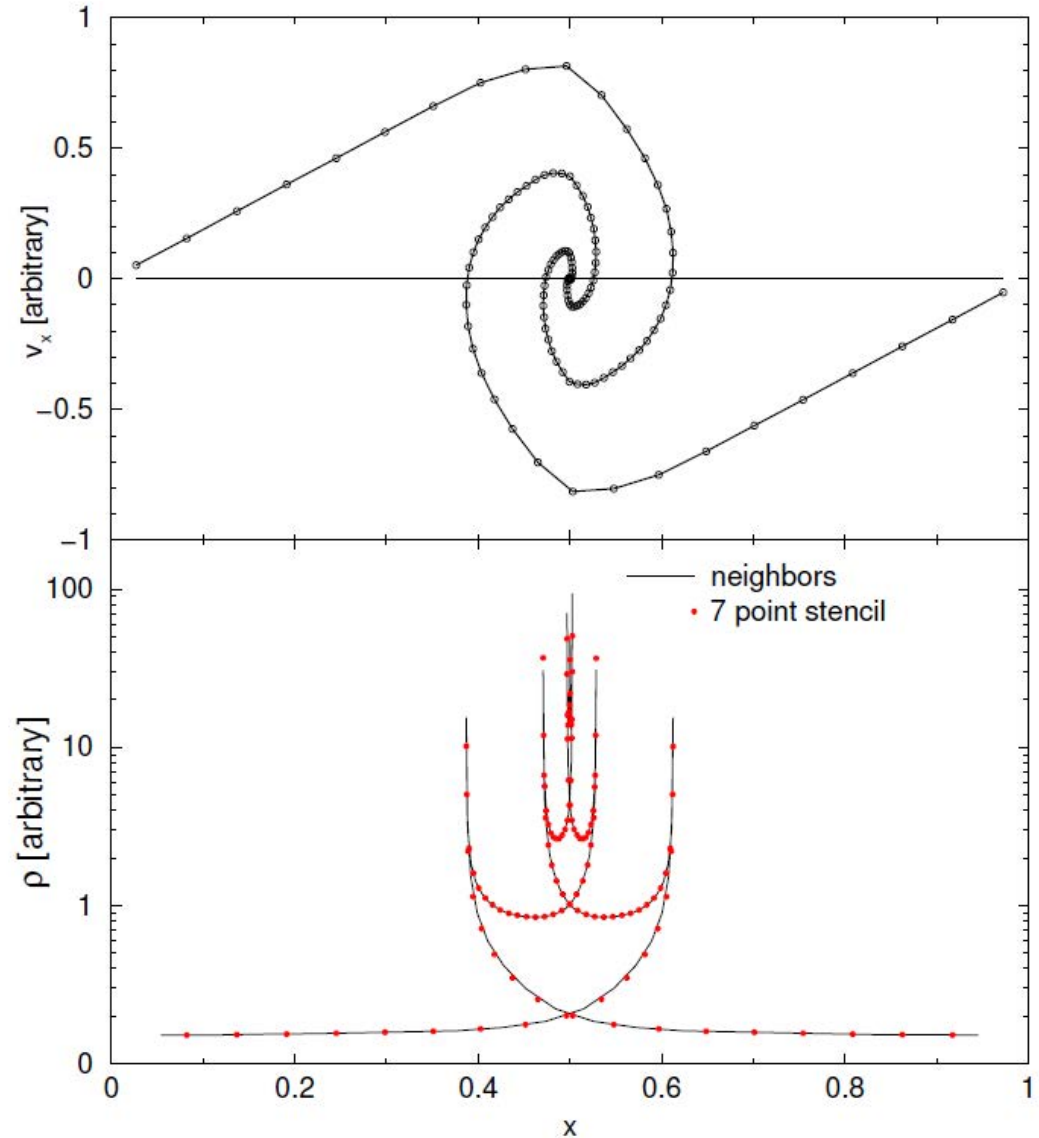
Phase space:

Velocity vs. Position



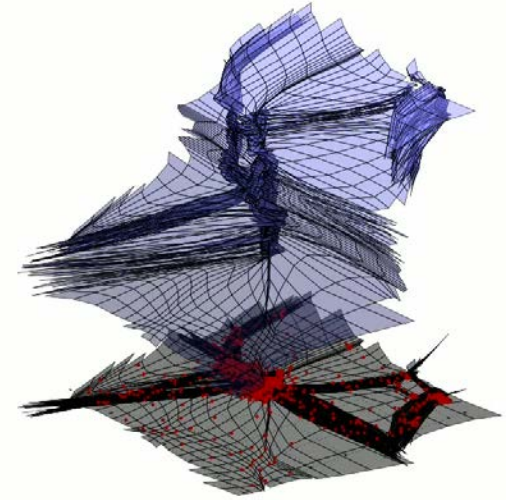
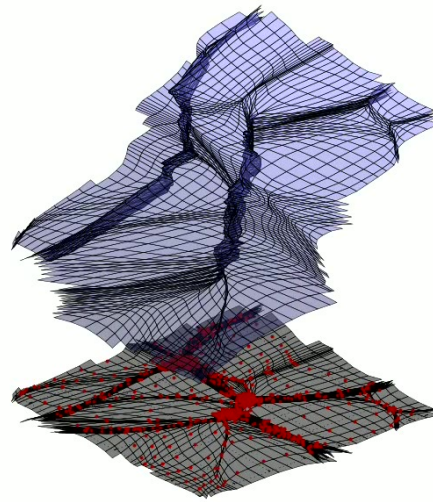
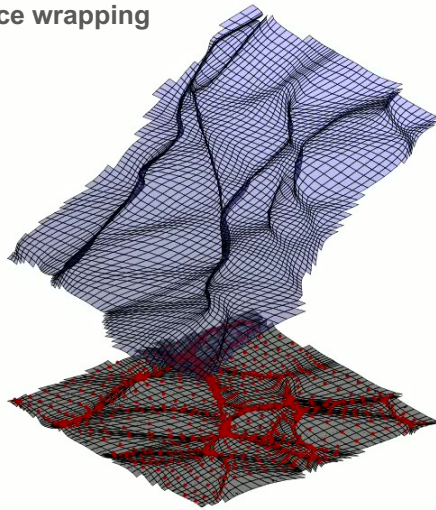
Density:

$$\rho(\vec{x}, t) = \int f(\vec{x}, \vec{v}, t) d\vec{v}$$



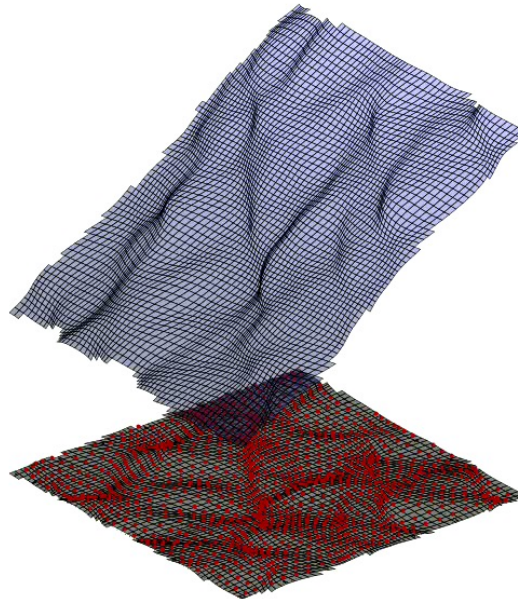
Phase-Space Sheet

Cosmic Web
phase-space wrapping



Dynamical Evolution:

folding the
phase-space sheet $\{q, x\}$

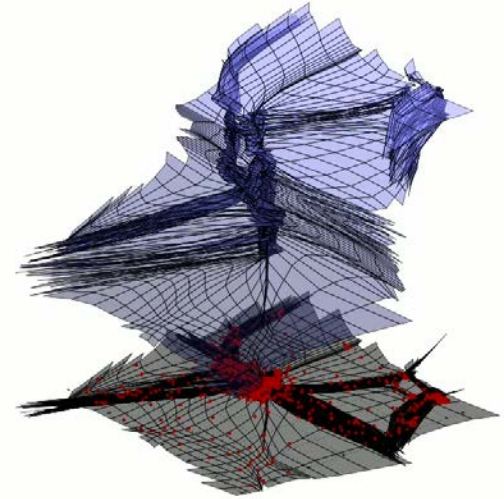
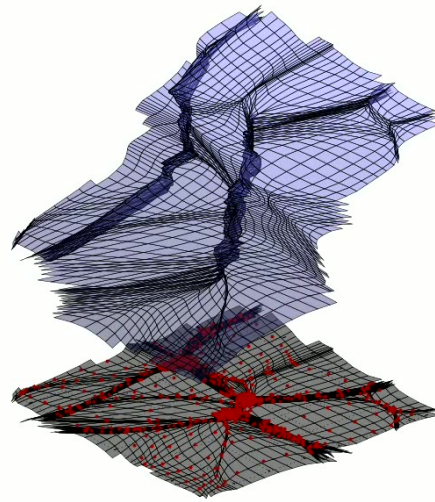
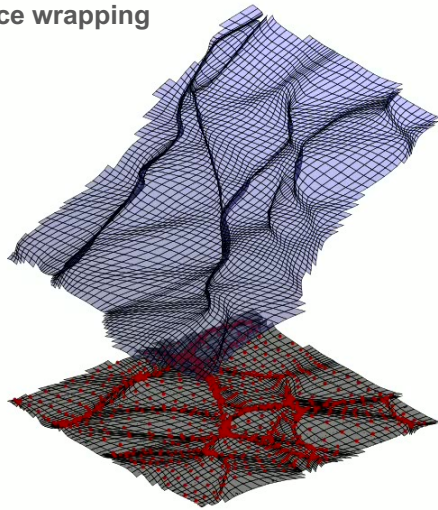


Lagrangian
coordinate

Eulerian plane
→

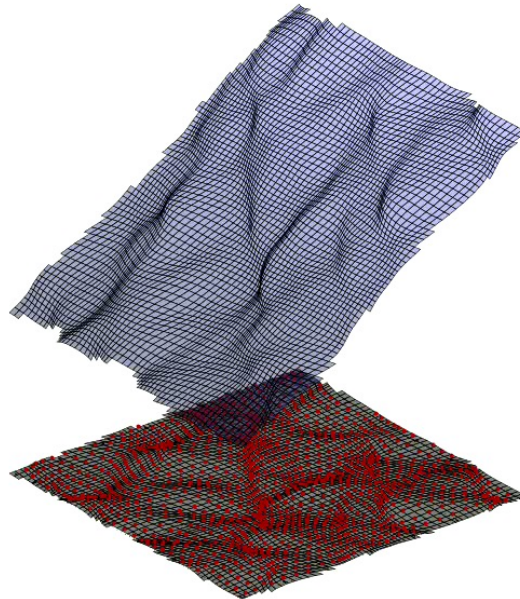
Phase-Space Sheet

Cosmic Web
phase-space wrapping

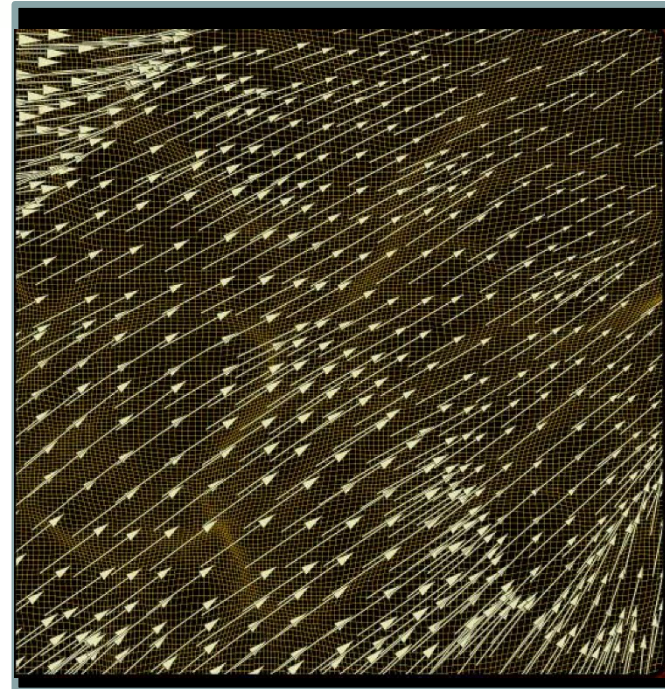


Dynamical Evolution:

folding the
phase-space sheet $\{q, x\}$



Hidding 2014

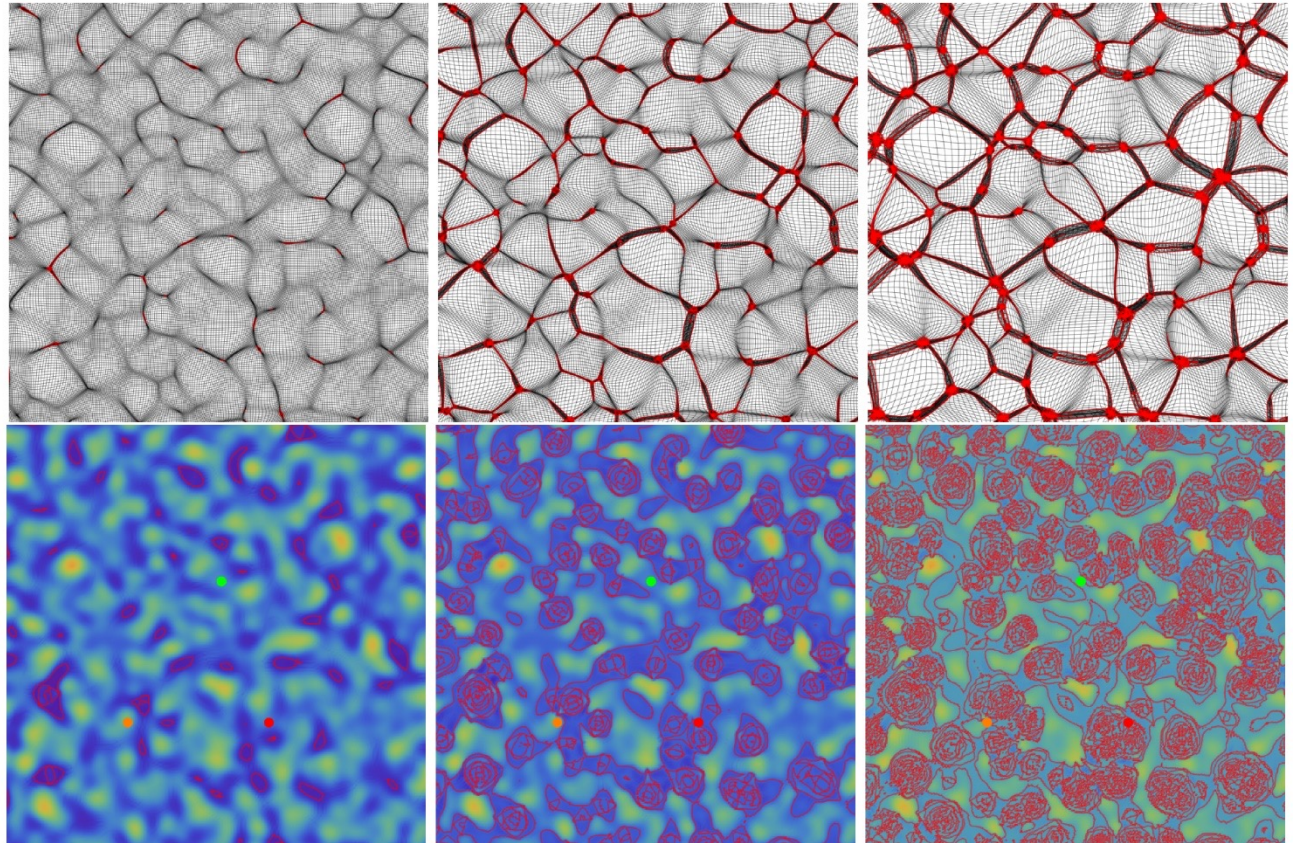


Cosmic Web Multistreaming

Translation towards
2D space:

**Evolution
Multistream region:**

- Eulerian space
- Lagrangian space
(mass elements)



Shandarin 2012
Abel, Hahn & Kaehler 2012
Falck, Neyrinck et al. 2012
Feldbrugge, Wilding, vdW 2022

Cosmic Web FlipFlop field

Translation towards
3D space:

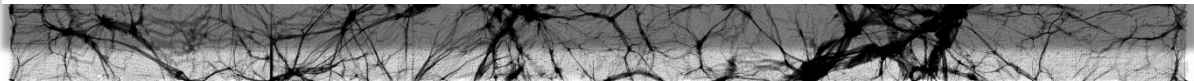
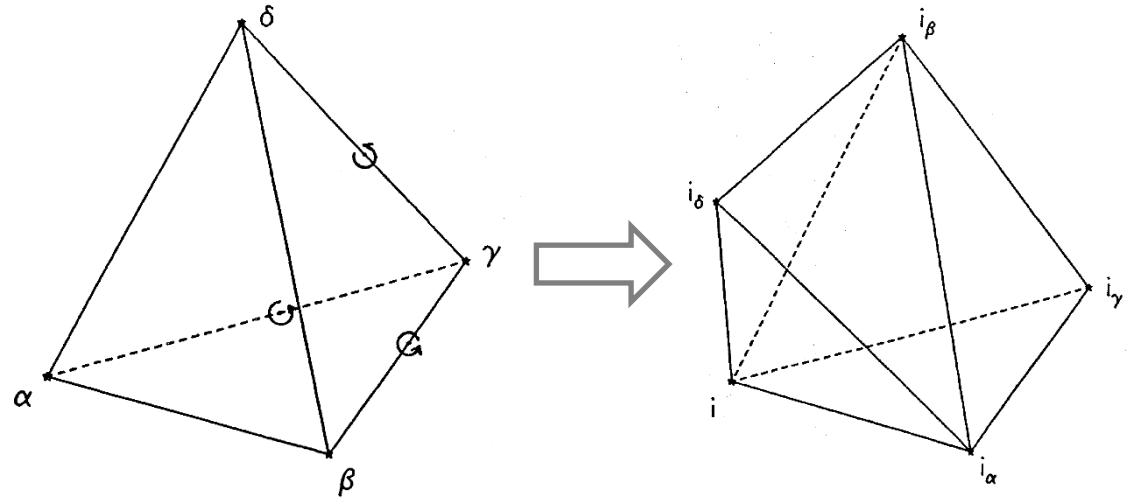
Density of
dark matter streams:

- # phase space folds

=

changing orientation
tetrahedra

Shandarin 2012
Feldbrugge et al. 2022b



Cosmic Web FlipFlop field

Translation towards
3D space:

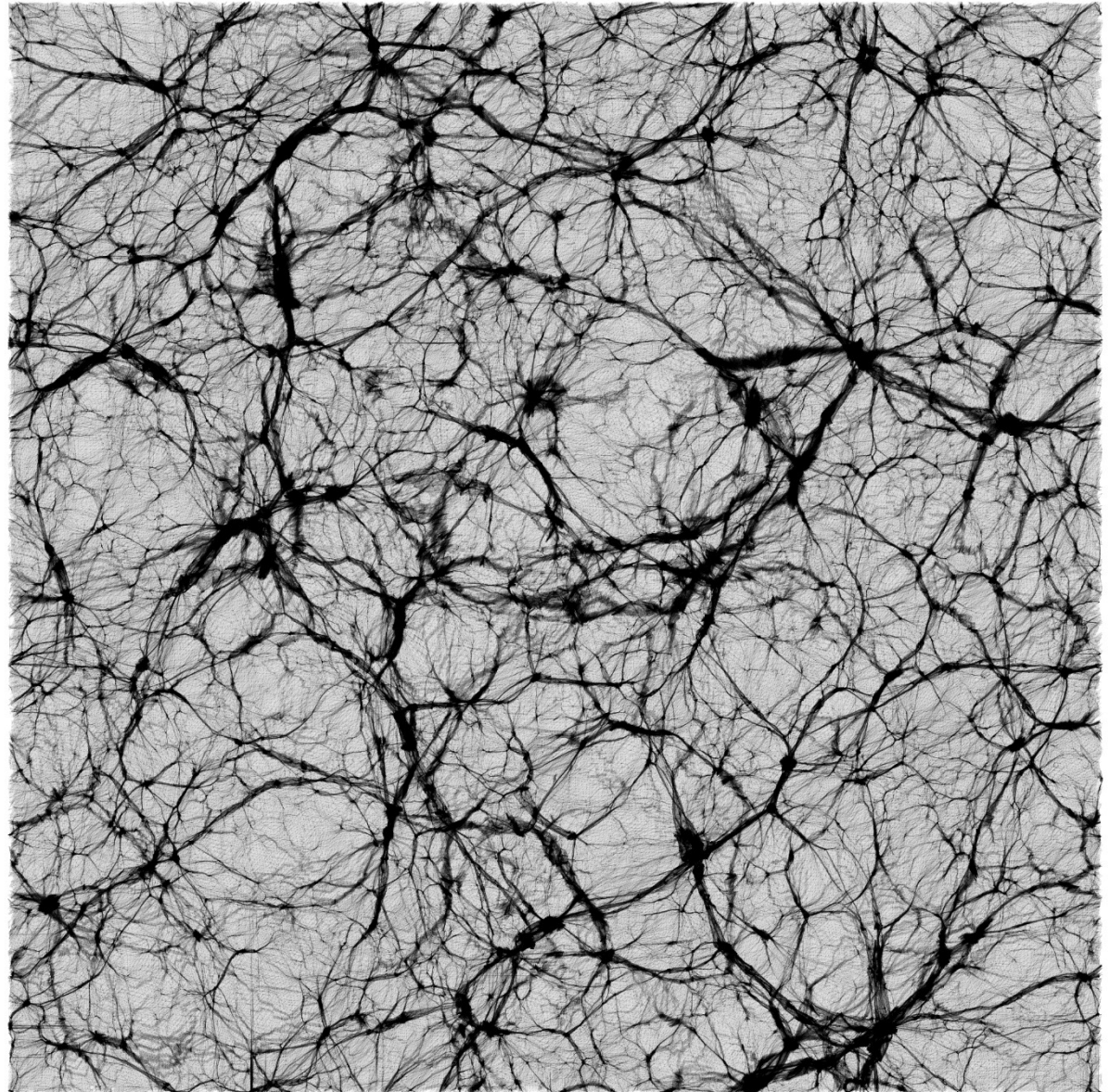
Density of
dark matter streams:

- # phase space folds

=

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tetrahedra

Shandarin 2012
Feldbrugge et al. 2022b



Cosmic Web FlipFlop field

Translation towards
3D space:

Density of
dark matter streams:

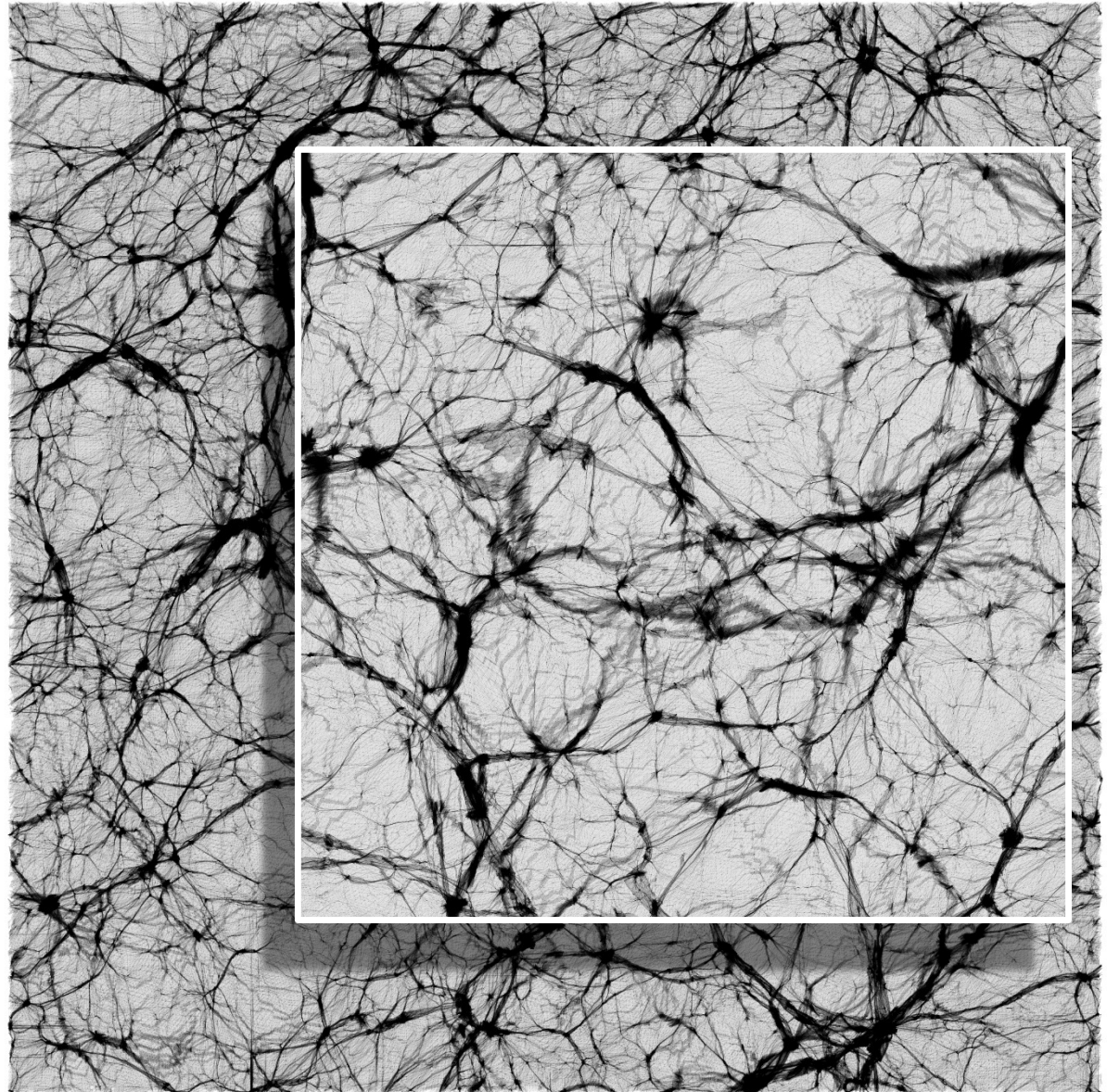
- # phase space folds

=

changing orientation
tetrahedra

Shandarin 2012

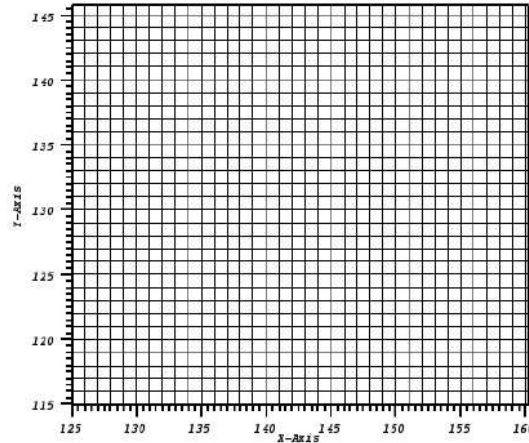
Feldbrugge et al. 2022b



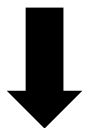
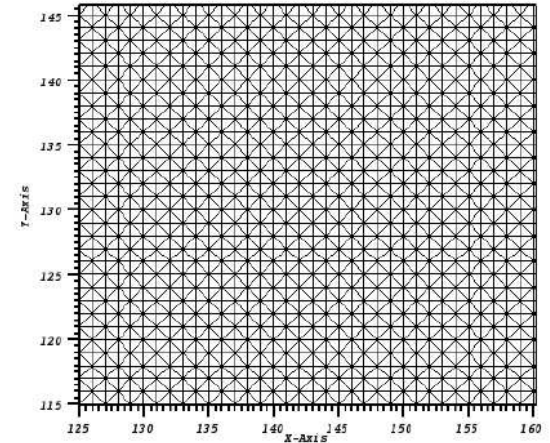
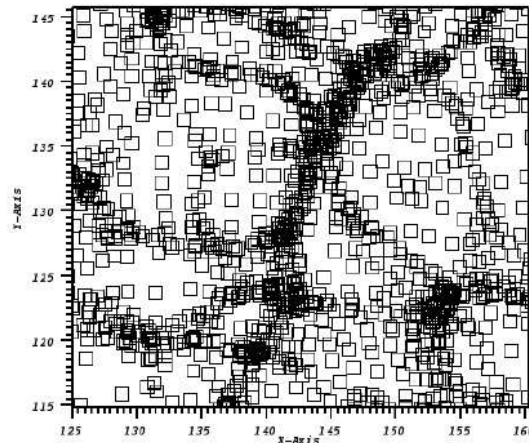
Tessellation Deformation & Phase Space Projection

Translation towards
Multi-D space:

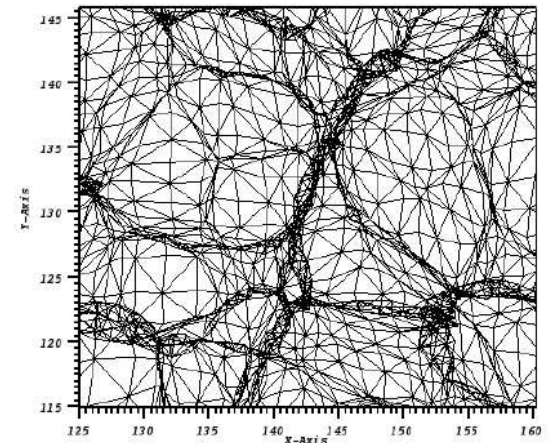
- Look at deformation of initial tessellation
- each tessellation cell represents matter cell
- evolution deforms cell
- once cells start to overlap, manifestation of different phase-space matter streams



particle
displacement



fluid element
deformation

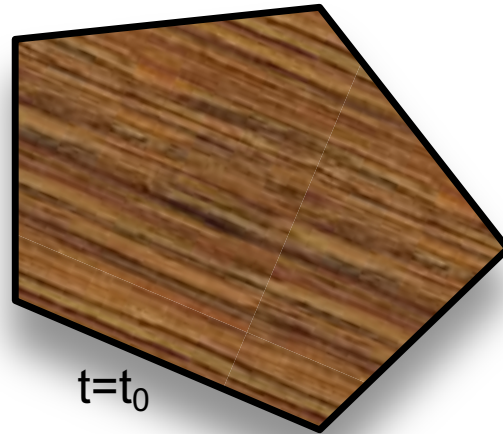


Tessellation Deformation & Phase Space Projection

Translation towards
Multi-D space:

- Look at deformation of initial tessellation
- each tessellation cell represents matter cell
- evolution deforms cell

Monostream
Density Evolution



$t=t_1$

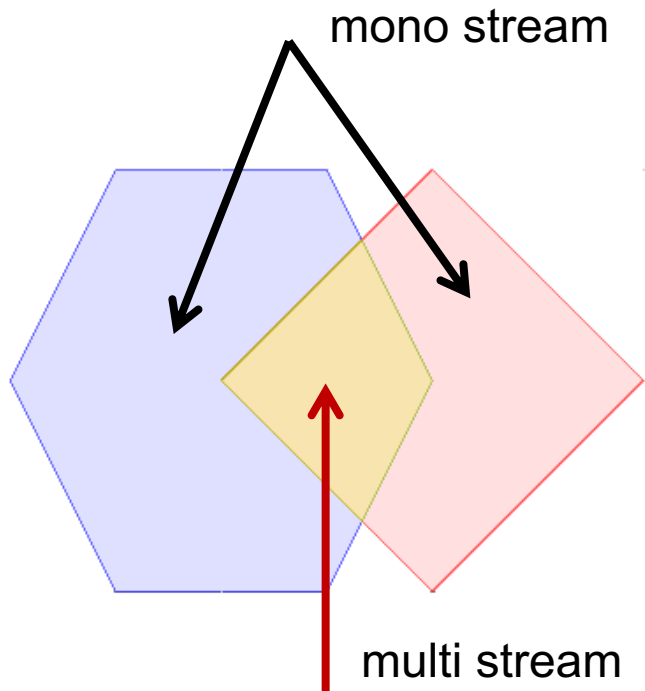
Conservation of mass
(continuity eqn.):

$$\rho(\vec{x}, t) = |J(\vec{x}, \vec{q})|^{-1} \rho(\vec{q}) = \left| \frac{\partial \vec{x}}{\partial \vec{q}} \right|^{-1} \rho(\vec{q})$$

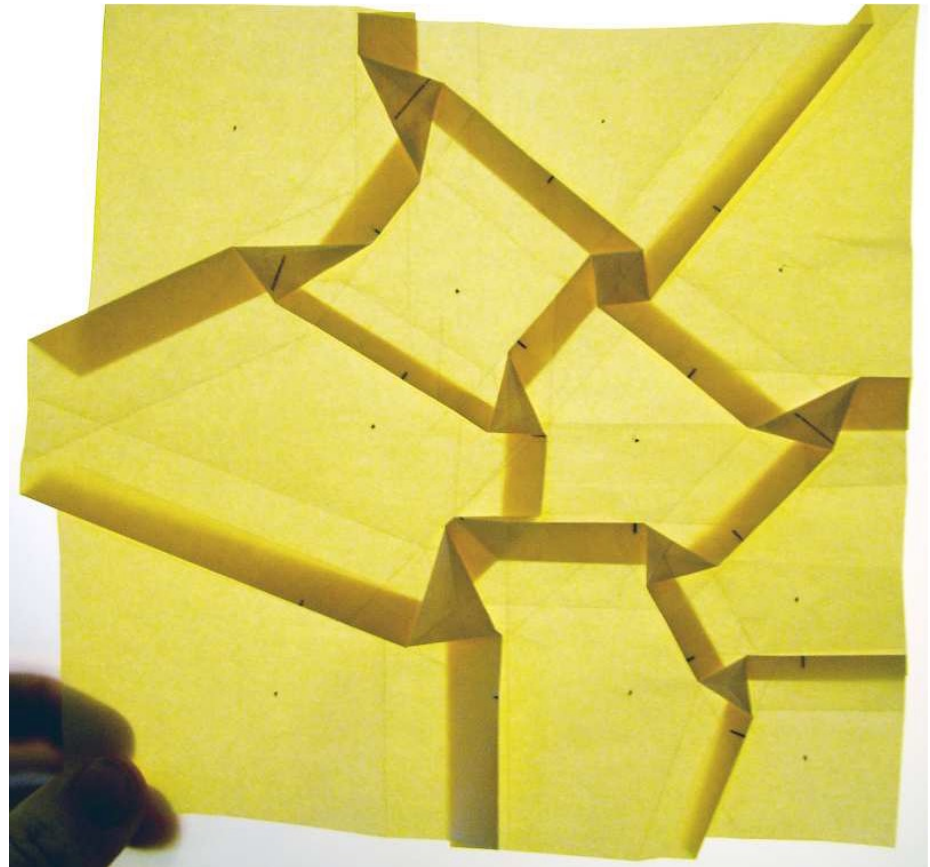


$$\rho(\vec{x}, t_1) = \frac{V_0}{V_1} \rho(\vec{q}, t_0)$$

(Cosmic) ORIGAMI

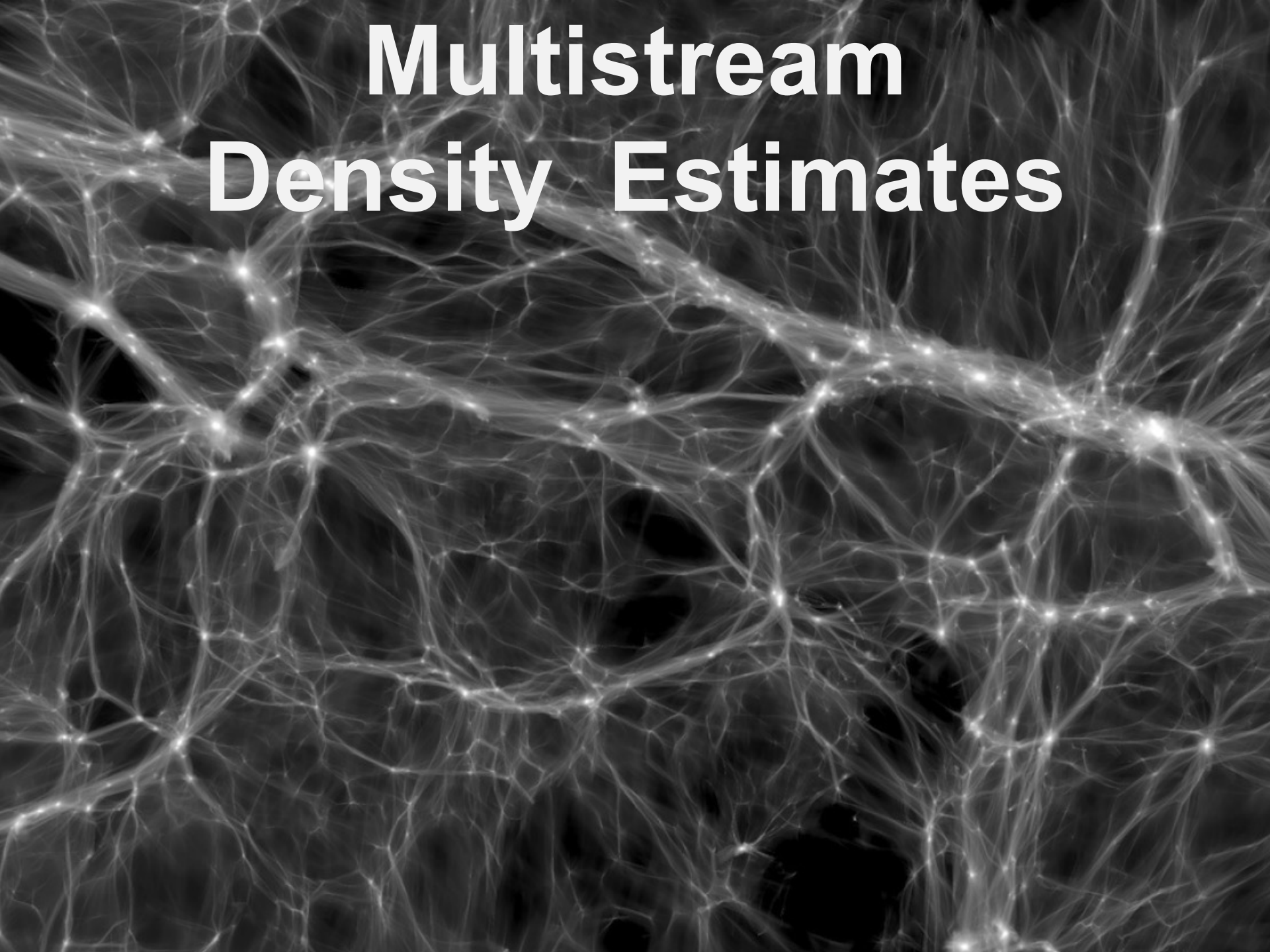


Evolution of dynamical system:
Phase-space folding – Cosmic Origami



$$\rho_{total}(\vec{x}, t_1) = \sum_i \frac{V_{0i}}{V_{1i}} \rho(\vec{q}_i, t_0)$$

Multistream Density Estimates



Cosmic Web Stream Density

Translation towards
Multi-D space:

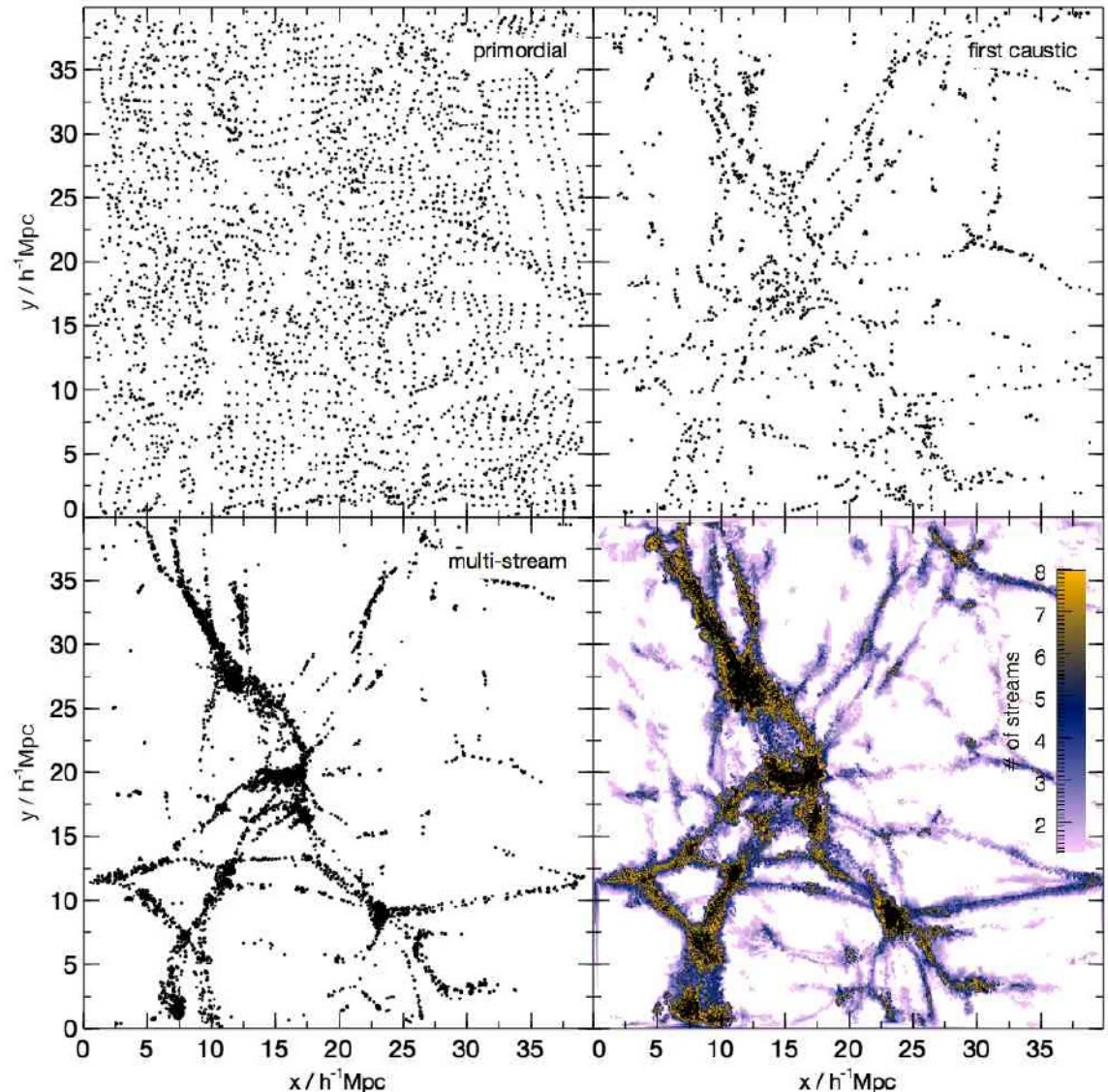
Density of
dark matter streams:

- # phase space folds

=

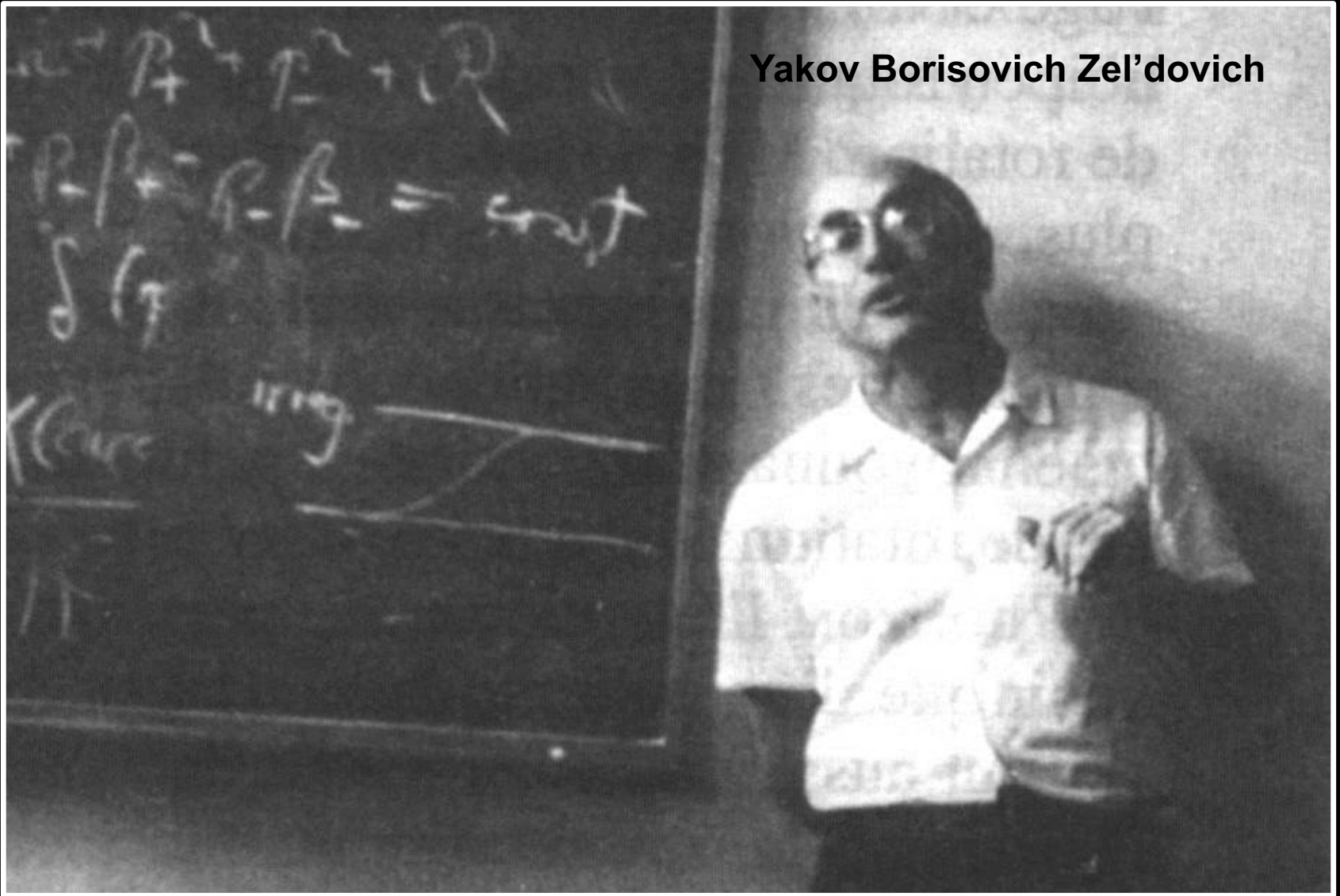
locally overlapping
tessellation cells

Shandarin 1992
Abel, Hahn & Kaehler 2012
Falck, Neyrinck et al. 2012



**1st order
Lagrangian Theory:
Zeldovich Formalism**

Yakov Borisovich Zel'dovich



Zel'dovich Approximation

1st order Lagrangian perturbation theory

$$\vec{x} = \vec{q} + D(t) \vec{u}(\vec{q})$$

$$\vec{u}(\vec{q}) = -\vec{\nabla} \Phi(\vec{q})$$

$$\Phi(\vec{q}) = \frac{2}{3Da^2H^2\Omega} \phi_{lin}(\vec{q})$$

$$\delta_{lin}(\vec{x}, t) = D(t) \delta(\vec{x}, t_0)$$

linear growth factor D

$$D(t) \approx H(t) \int \frac{dt}{a^2 H^2(t)}$$

$$\vec{u}(\vec{q}) = \frac{d\vec{x}}{dD};$$

$$\vec{v}(\vec{q}) = HDf(\Omega) \vec{u}(\vec{q})$$

$$f(\Omega) = \frac{d \log D}{d \log a}$$

linear growth rate f

Zel'dovich Approximation

$$\vec{x} = \vec{q} + D(t) \vec{u}(\vec{q})$$

$$\vec{u}(\vec{q}) = -\vec{\nabla} \Phi(\vec{q})$$

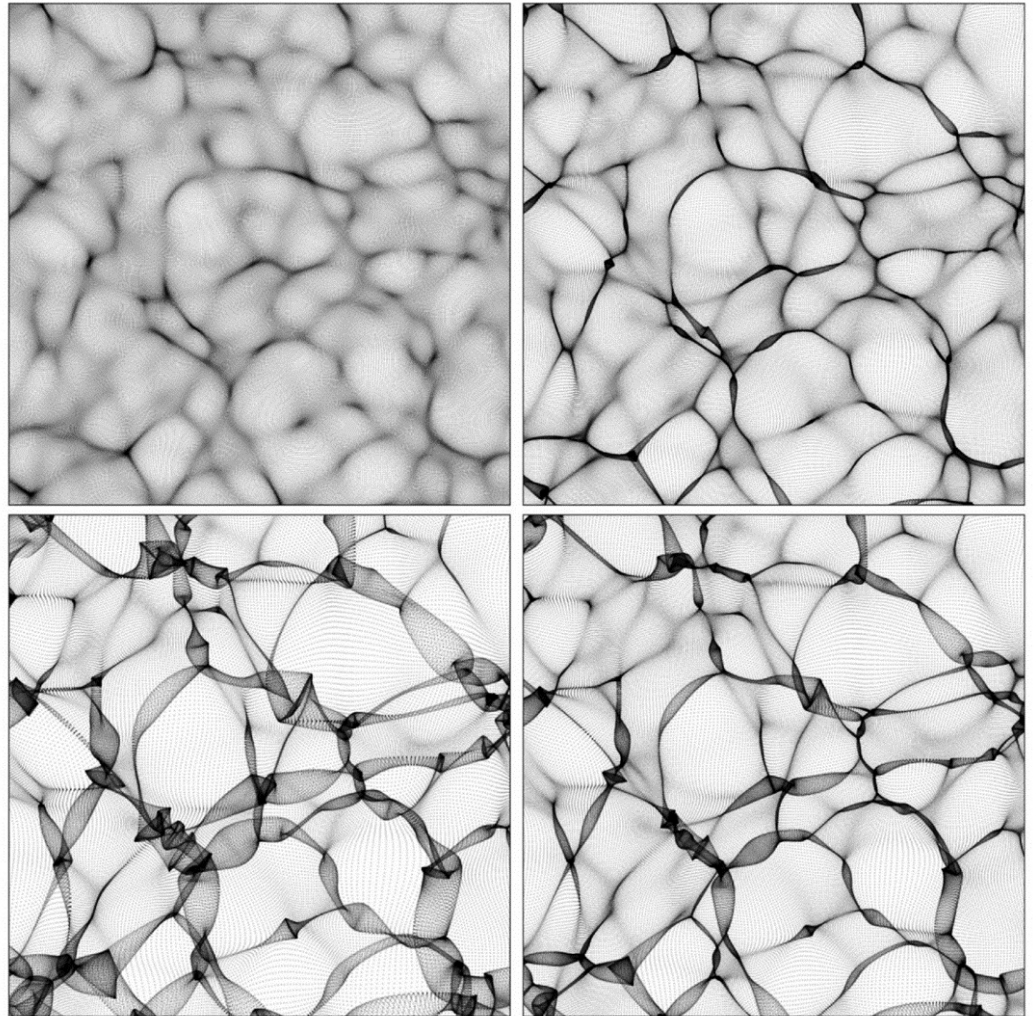
$$\Phi(\vec{q}) = \frac{2}{3Da^2 H^2 \Omega} \phi_{lin}(\vec{q})$$

Zel'dovich Approximation

$$\vec{x} = \vec{q} + D(t) \vec{u}(\vec{q})$$

$$\vec{u}(\vec{q}) = -\vec{\nabla} \Phi(\vec{q})$$

$$\Phi(\vec{q}) = \frac{2}{3Da^2 H^2 \Omega} \phi_{lin}(\vec{q})$$

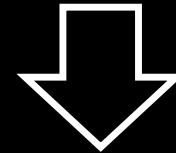


Zel'dovich Approximation: Deformation

$$\vec{x} = \vec{q} + D(t) \vec{u}(\vec{q})$$

$$\vec{u}(\vec{q}) = -\vec{\nabla} \Phi(\vec{q})$$

$$d_{ij} = -\frac{\partial u_i}{\partial q_j}$$



$$\rho(\vec{q}, t) = \frac{\rho_u(t)}{(1 - D(t)\lambda_1(\vec{q}))(1 - D(t)\lambda_2(\vec{q}))(1 - D(t)\lambda_3(\vec{q}))}$$

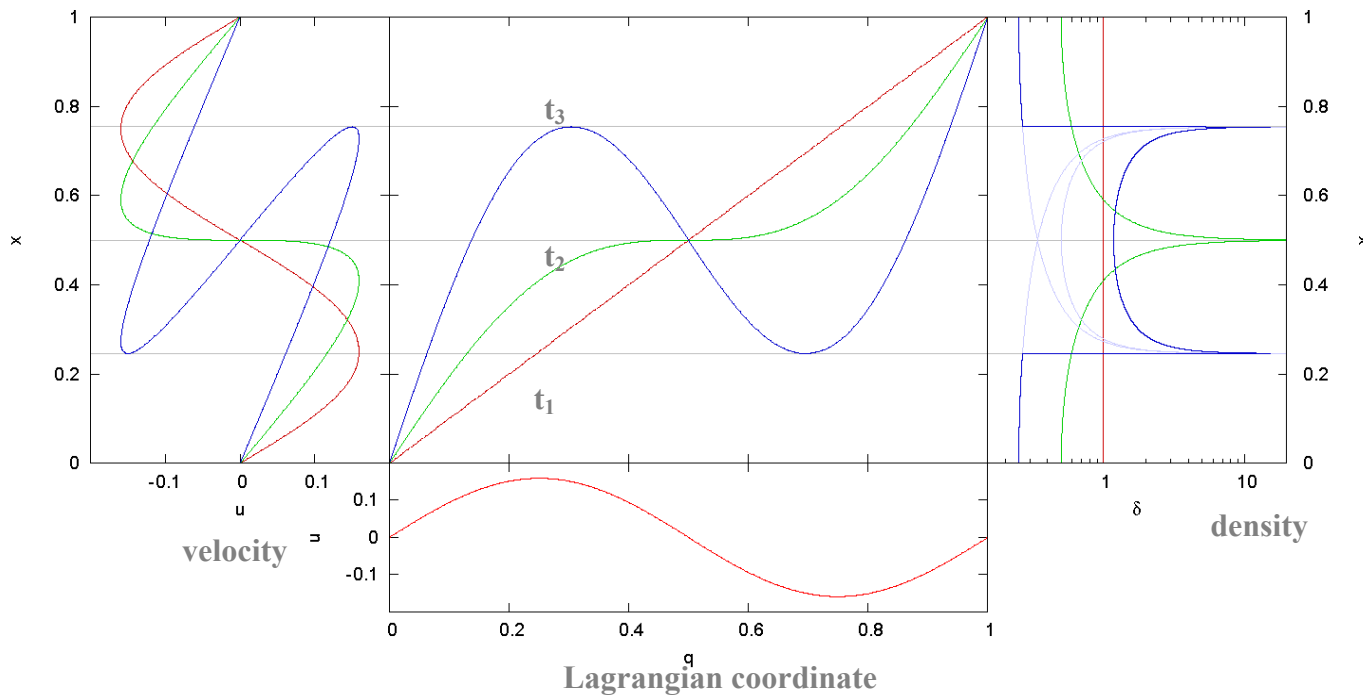
structure of the cosmic web determined by the
spatial field of eigenvalues

$$\lambda_1, \lambda_2, \lambda_3$$

Zeldovich Formalism:

Singularities

Eulerian coordinate



$$\vec{x}(\vec{q}, t) = \vec{q} - D(t) \vec{\nabla} \Phi(\vec{q}) \quad \Rightarrow \quad d_{ij} = \frac{\partial^2 \Phi}{\partial q_i \partial q_j} : \lambda_1, \lambda_2, \lambda_3$$

$$\rho(\vec{q}, t) = \frac{\rho_u(t)}{(1 - D(t)\lambda_1(\vec{q}))(1 - D(t)\lambda_2(\vec{q}))(1 - D(t)\lambda_3(\vec{q}))}$$

Zel'dovich Morphology

$$\rho(\vec{q}, t) = \frac{\rho_u(t)}{(1 - D(t)\lambda_1(\vec{q}))(1 - D(t)\lambda_2(\vec{q}))(1 - D(t)\lambda_3(\vec{q}))}$$

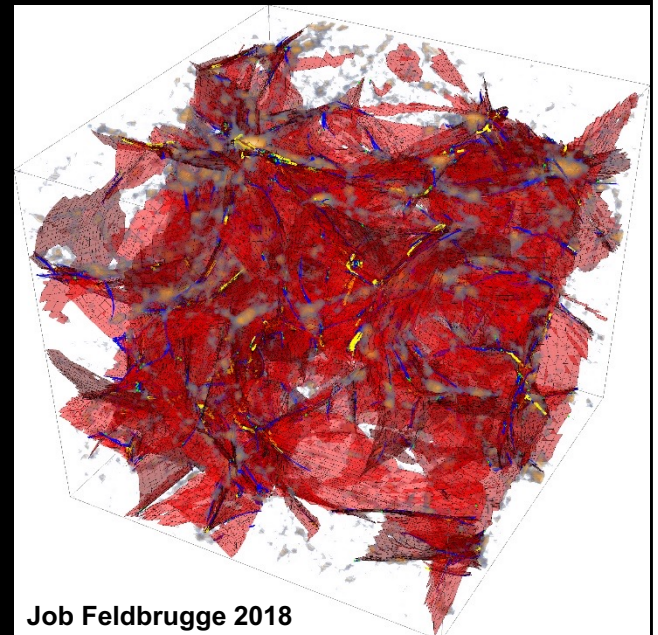
$$\lambda_1, \lambda_2, \lambda_3$$

$$\lambda_1 > \lambda_2 > \lambda_3$$

Structure of the cosmic web determined by the spatial field of eigenvalues:

Sequence of formation stages:

- λ_1 - collapse along first axis:
formation of walls/sheets/pancakes
- λ_2 - collapse along 2 axes:
formation of elongated filaments
- λ_3 - possibly – if $\lambda_3 > 0$ – collapse along all three axes, into a fully collapsed clump/node



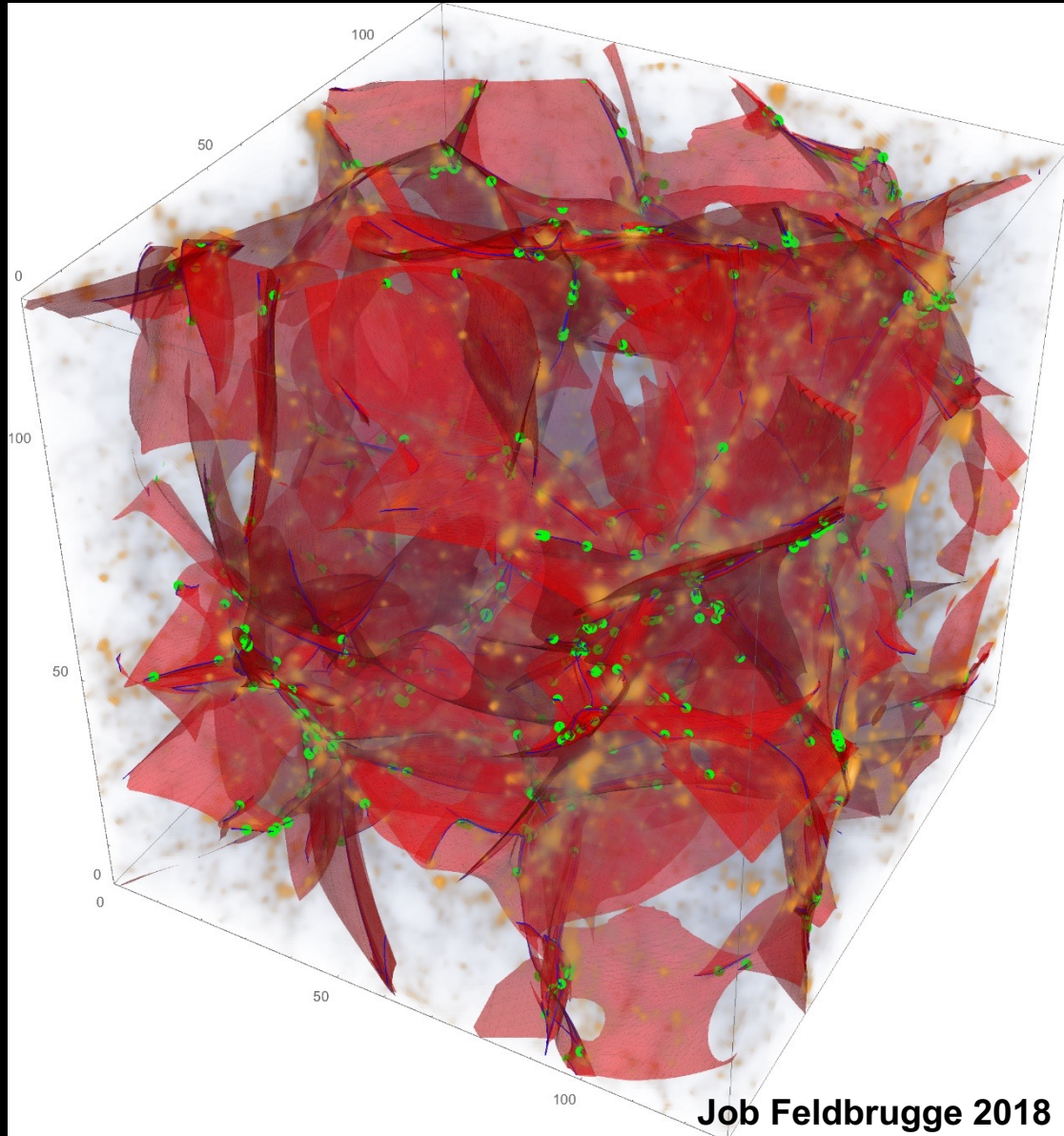
Job Feldbrugge 2018

Zel'dovich Cosmic Web

It is no exaggeration to state that Zeldovich (1970) predicted the existence of the Cosmic Web !

Sequence of formation stages:

- λ_1 - collapse along first axis:
formation of walls/sheets/pancakes
- λ_2 - collapse along 2 axes:
formation of elongated filaments
- λ_3 - possibly – if $\lambda_3 > 0$ – collapse along all three axes,
into a fully collapsed clump/node



Job Feldbrugge 2018

Yakov Borisovich Zel'dovich



Zel'dovich Dynamics

1st order Lagrangian perturbation theory

$$\vec{x} = \vec{q} + D(t)\vec{u}(\vec{q})$$

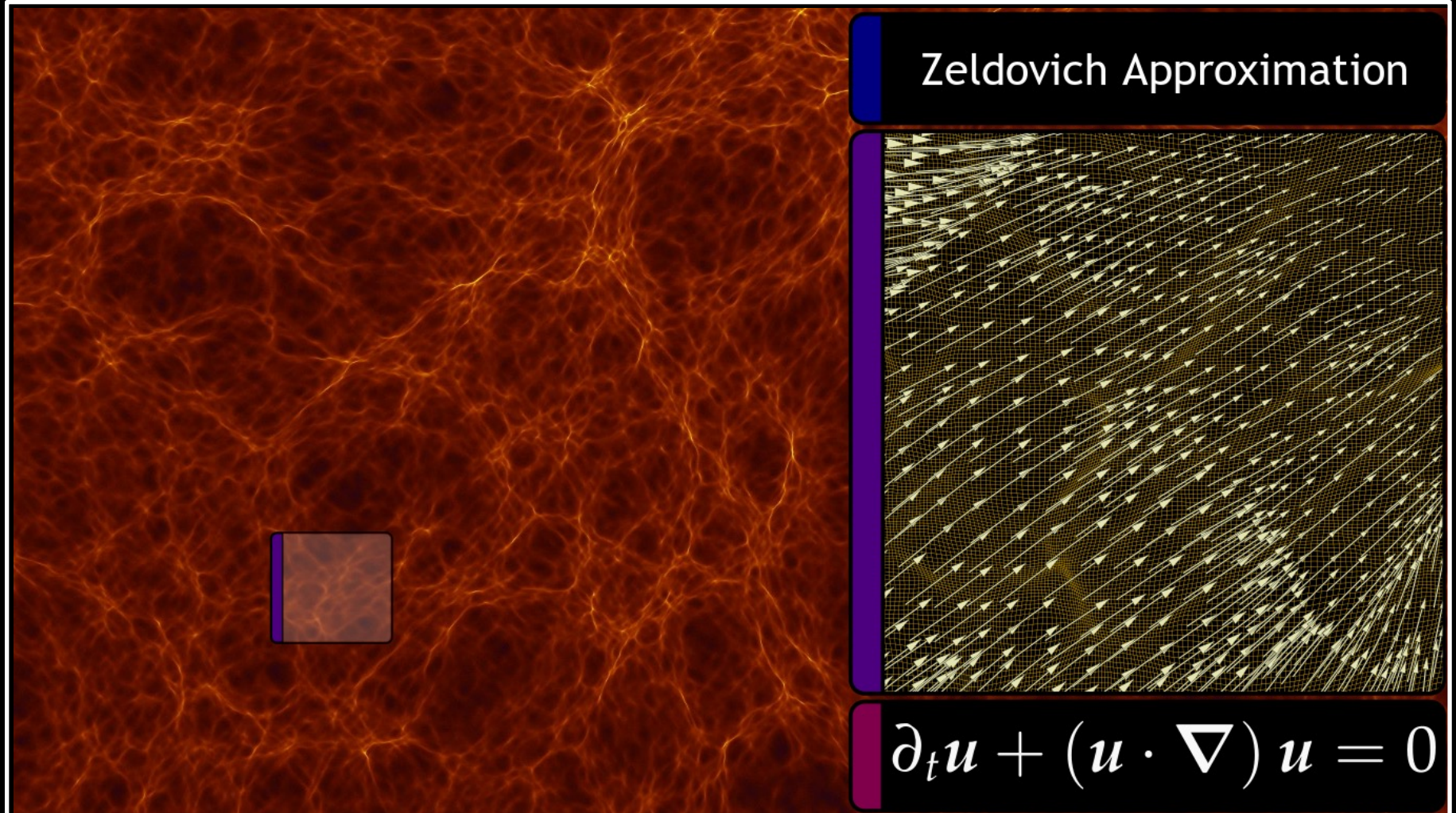
$$\vec{u}(\vec{q}) = -\vec{\nabla}\Phi(\vec{q})$$

Euler equation:

force-free flow

$$\frac{\partial \vec{u}}{\partial D} + (\vec{u} \cdot \nabla) \vec{u} = 0$$

Zel'dovich Dynamics



Zeldovich Dynamics

By rewriting the Euler equation (in comoving coordinates), we may easily understand dynamical nature of the Zeldovich approximation:

$$\frac{\partial \vec{v}}{\partial t} + \frac{\dot{a}}{a} \vec{v} + \frac{1}{a} (\vec{v} \cdot \vec{\nabla}) \vec{v} = -\frac{1}{a} \vec{\nabla} \phi$$

Define velocity \vec{u} ,
wrt linear growth factor $D(t)$:

$$\vec{u} = \frac{d\vec{x}}{dD} = \frac{\vec{v}}{a\dot{D}}$$

Zeldovich Dynamics

Following some algebraic manipulations, one arrives at the equivalent Euler equation for the normalized velocity \vec{u} :

$$\frac{\partial \vec{u}}{\partial D} + (\vec{u} \cdot \vec{\nabla}) \vec{u} = -\vec{\nabla} \left(\frac{3\Omega}{2f^2 D} \phi_v + \frac{\phi}{a^2 \dot{D}^2} \right) = -\vec{\nabla} V$$

With velocity potential ϕ_v :

$$\vec{u} = -\vec{\nabla} \phi_v$$

and effective potential V :

$$V = \frac{3\Omega}{2f^2 D} (\phi_v + \theta)$$

and scaled
gravitational potential θ :

$$\theta = \frac{2\phi}{3\Omega a^2 D H^2}$$

Effective & Scaled Potentials

For the Zeldovich approximation we may easily see that the effective potential $V=0$:

$$V = \frac{3\Omega}{2f^2 D} (\phi_v + \theta) = 0$$

For the Zeldovich approximation:

$$\vec{x} = \vec{q} - D(t) \vec{\nabla} \Psi(\vec{q})$$

with:

$$\Psi(\vec{q}) = \frac{2}{3Da^2 H^2 \Omega} \phi(\vec{x}, t)$$

so that the scaled
gravitational potential θ :

$$\theta = \frac{2\phi}{3\Omega a^2 D H^2} = \Psi(\vec{q})$$

The velocity potential ϕ_v we may infer from the velocity corresponding to the Zeldovich approximation:

$$\vec{v} = \dot{\vec{x}} = -aDH f(\Omega) \vec{\nabla} \Psi(\vec{q})$$

$$\vec{u} = \vec{\nabla} \phi_v = \frac{\vec{v}}{a\dot{D}} = -\frac{aDH}{a\dot{D}} f(\Omega) \vec{\nabla} \Psi(\vec{q}) = -\vec{\nabla} \Psi(\vec{q})$$

from which we see that

$$\phi_v = -\Psi(\vec{q})$$

Hence, for the Zeldovich approximation: $\phi_v + \theta = 0 \quad \Rightarrow \quad V = 0$

Zel'dovich ++:

Adhesion Formalism

&

Hierarchical Cosmic Web Dynamics

Zeldovich-Adhesion

We saw that dynamically, the Zeldovich approximation corresponds to a force-free propagation, as evidenced by the Euler equation for the normalized velocity \vec{u} :

$$\frac{\partial \vec{u}}{\partial D} + \left(\vec{u} \cdot \vec{\nabla} \right) \vec{u} = -\vec{\nabla} V = 0$$

The force-free nature of the Zeldovich approximation leads to the ballistic motion, which once a mass element enters a multi-stream nonlinear region ignores the dominant self-gravitational terms, ie. the evolving gravitational potential of high-density structures (such as walls, filaments and clumps).

The adhesion approximation augments this with a (really) artificial term – a non-gravitational term – in terms of a viscosity term (as we know from the Navier-Stokes equation):

$$\frac{\partial \vec{u}}{\partial D} + \left(\vec{u} \cdot \vec{\nabla} \right) \vec{u} = \nu \nabla^2 \vec{u}$$

Zeldovich-Adhesion

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$$\frac{\partial \vec{u}}{\partial D} + (\vec{u} \cdot \vec{\nabla}) \vec{u} = \nu \nabla^2 \vec{u}$$

This equation, the Navier-Stokes equation for a pressureless medium, goes by the name of

Burger's Equation

after the famous hydrodynamicist. It is one of the few equations that can be fully solved analytically.

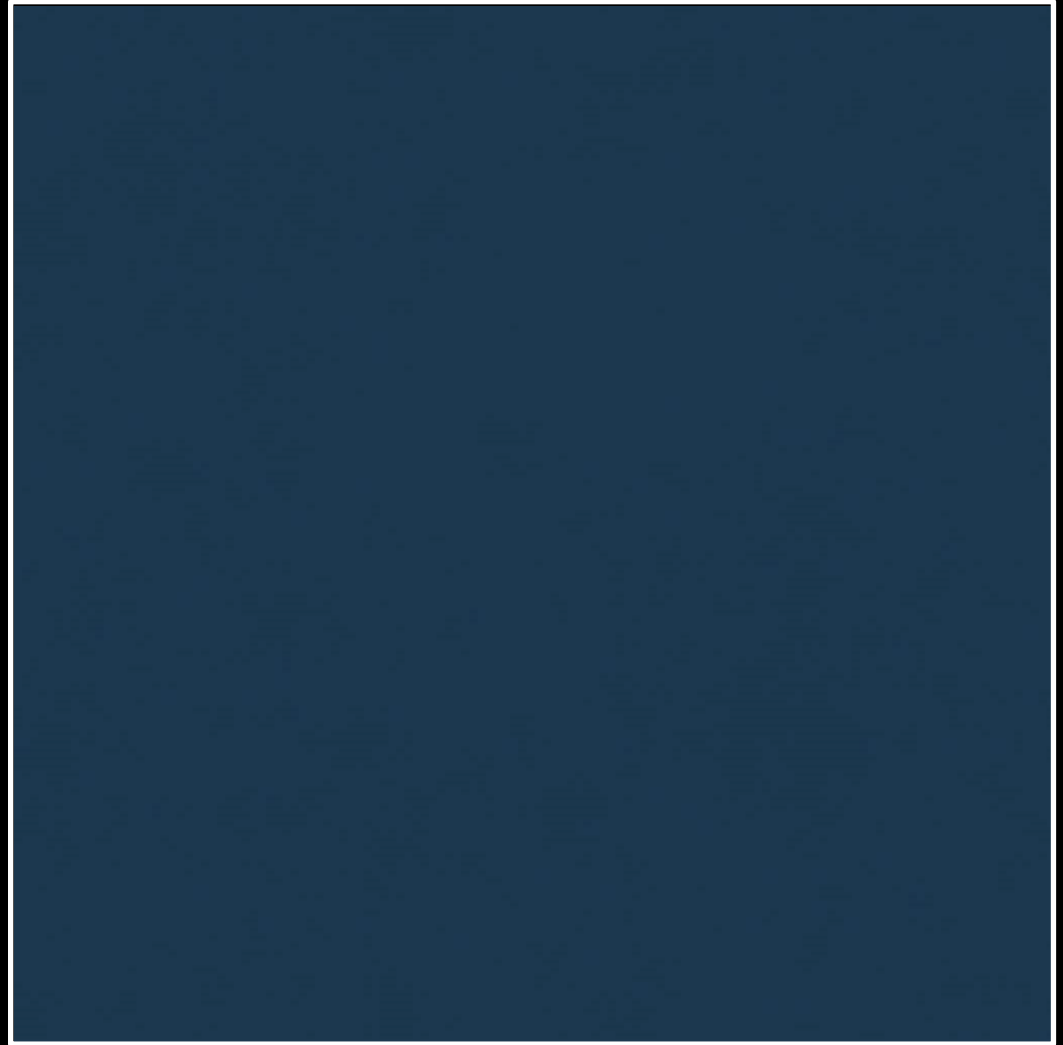
The viscosity term here is fully artificial, tries to emulate “selfgravity”, and has nothing to do with the physical viscosity we know from hydrodynamics. Basically, it functions as a friction term.

In its cosmological context, you only want to invoke it close to the emerging multistream regions, so that you take the asymptotic “inviscid” limit, $\nu \rightarrow 0$

Adhesion Approximation

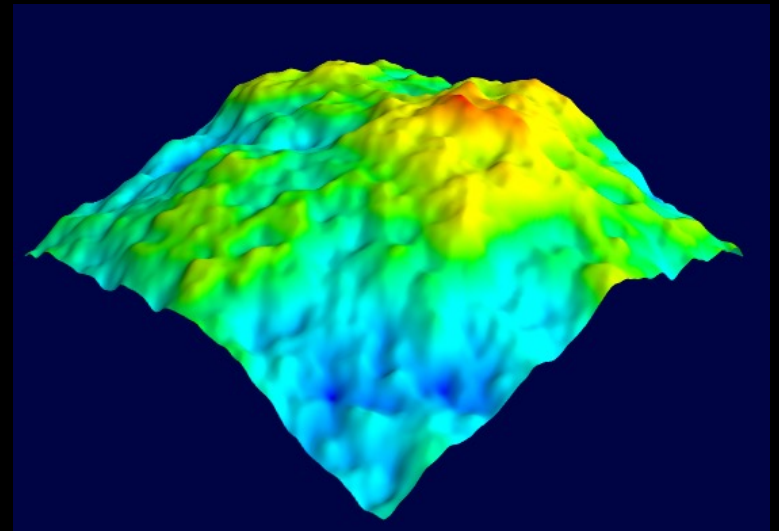
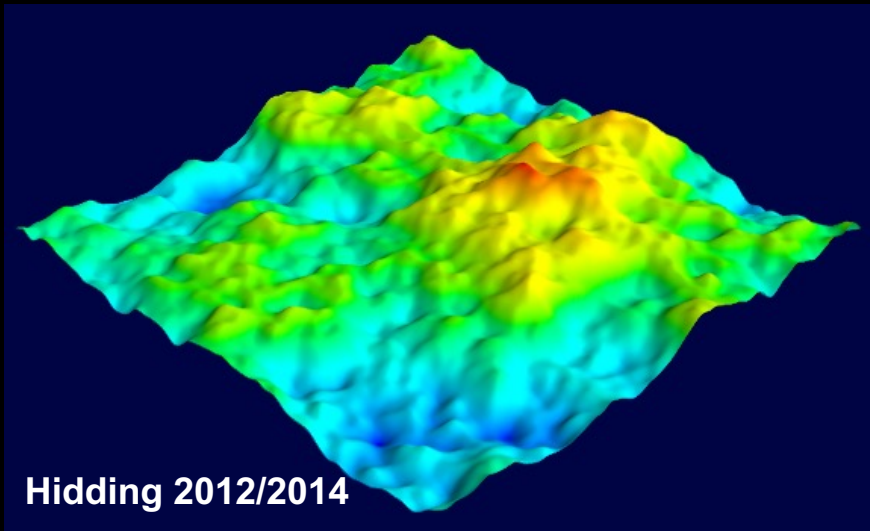
Gurbatov, Saichev & Shandarin 1987

Hidding 2012



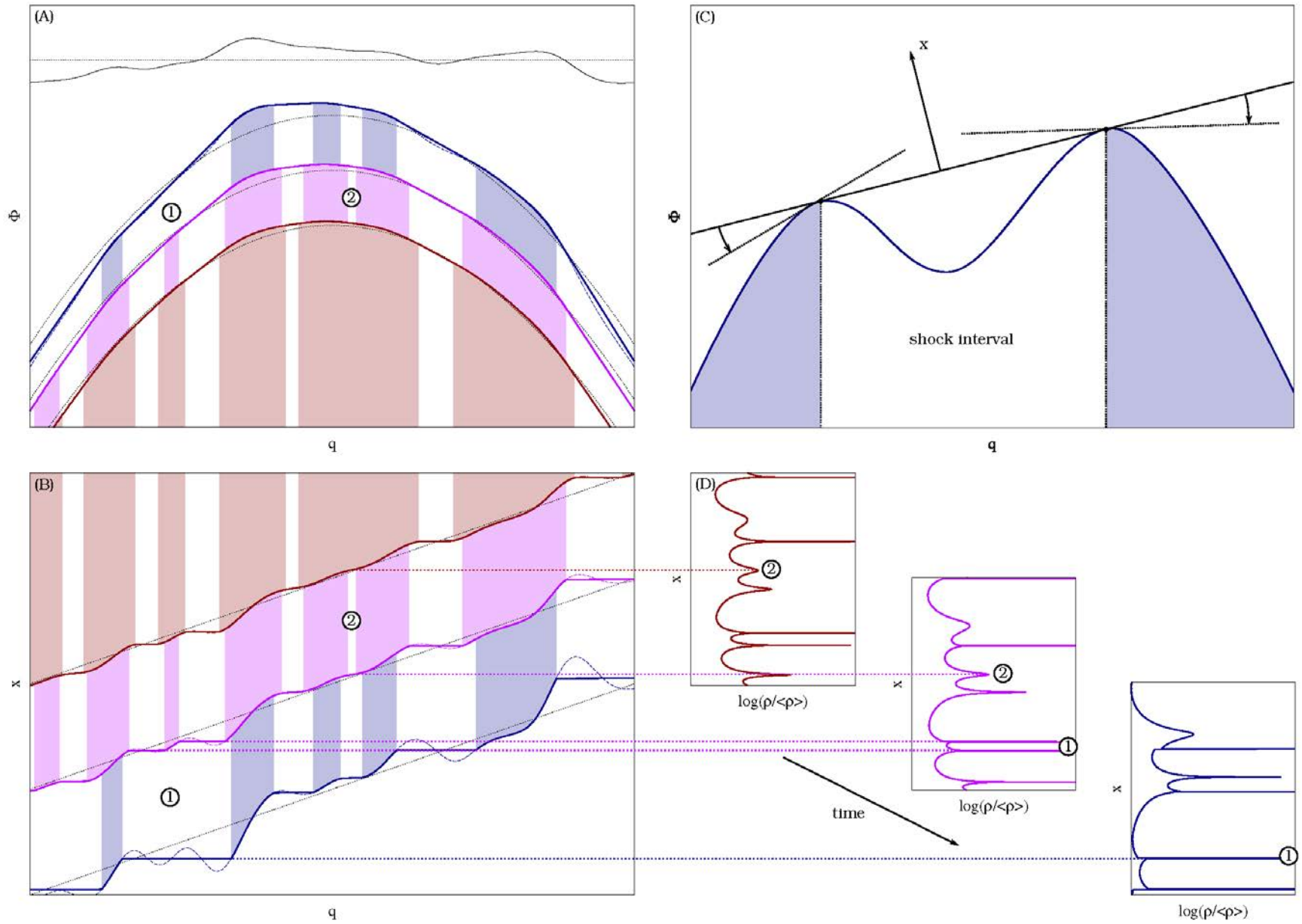
Burger's Equation: Hopf Solution

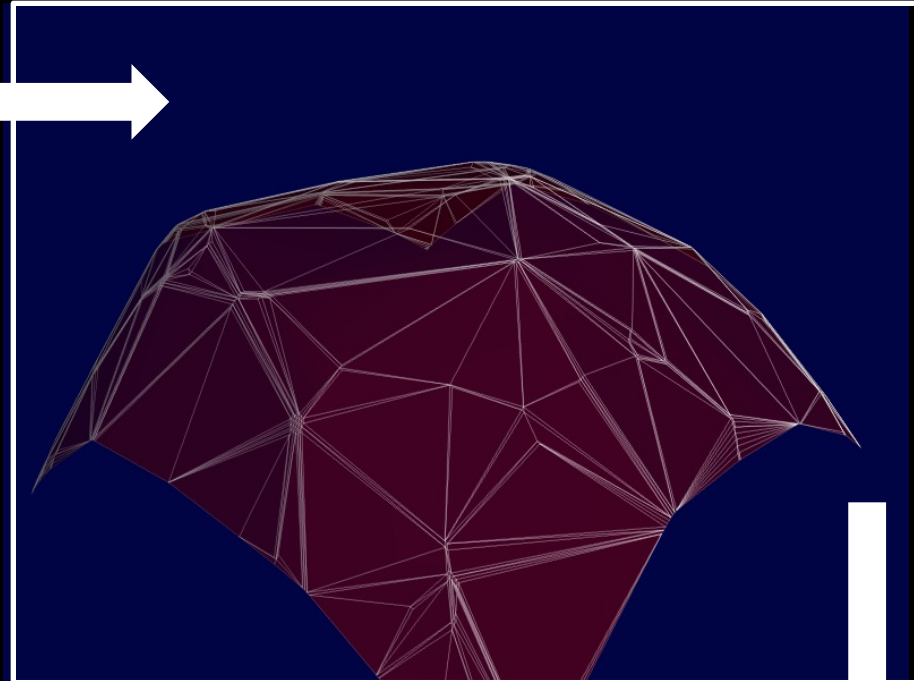
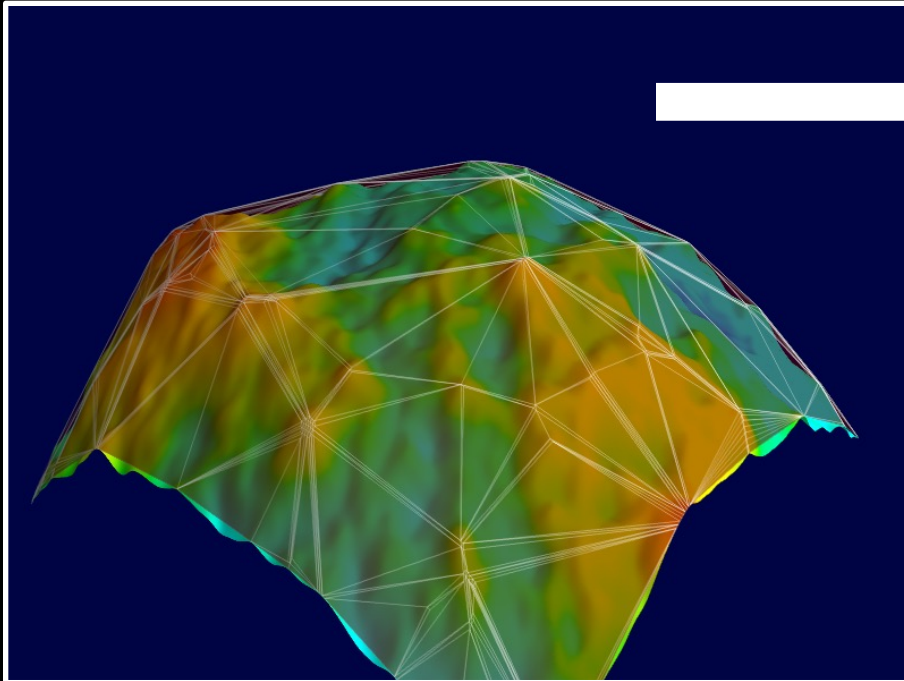
$$\frac{\partial \vec{u}}{\partial t} + (\vec{u} \cdot \vec{\nabla}) \vec{u} = \nu \nabla^2 \vec{u}$$



$$\Phi(\vec{x}, t) + \frac{x^2}{2} = \max_q \left[\left(t \Phi_0(q) - \frac{q^2}{2} \right) + \vec{x} \cdot \vec{q} \right]$$

Burger's Equation: Hopf Solution



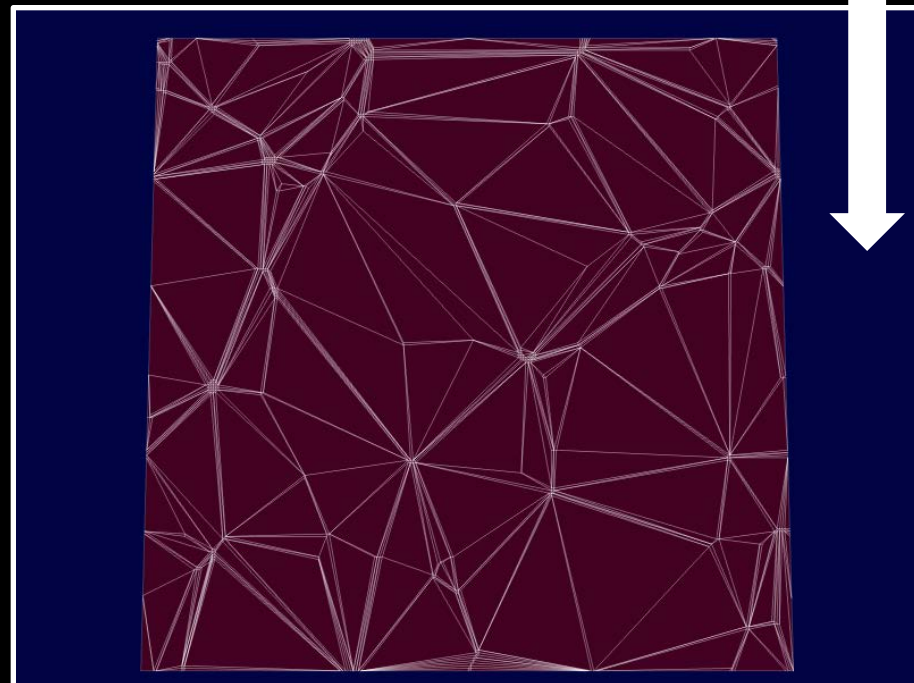


Hidding 2012/2014

Convex Hull
quadratically lifted potential field

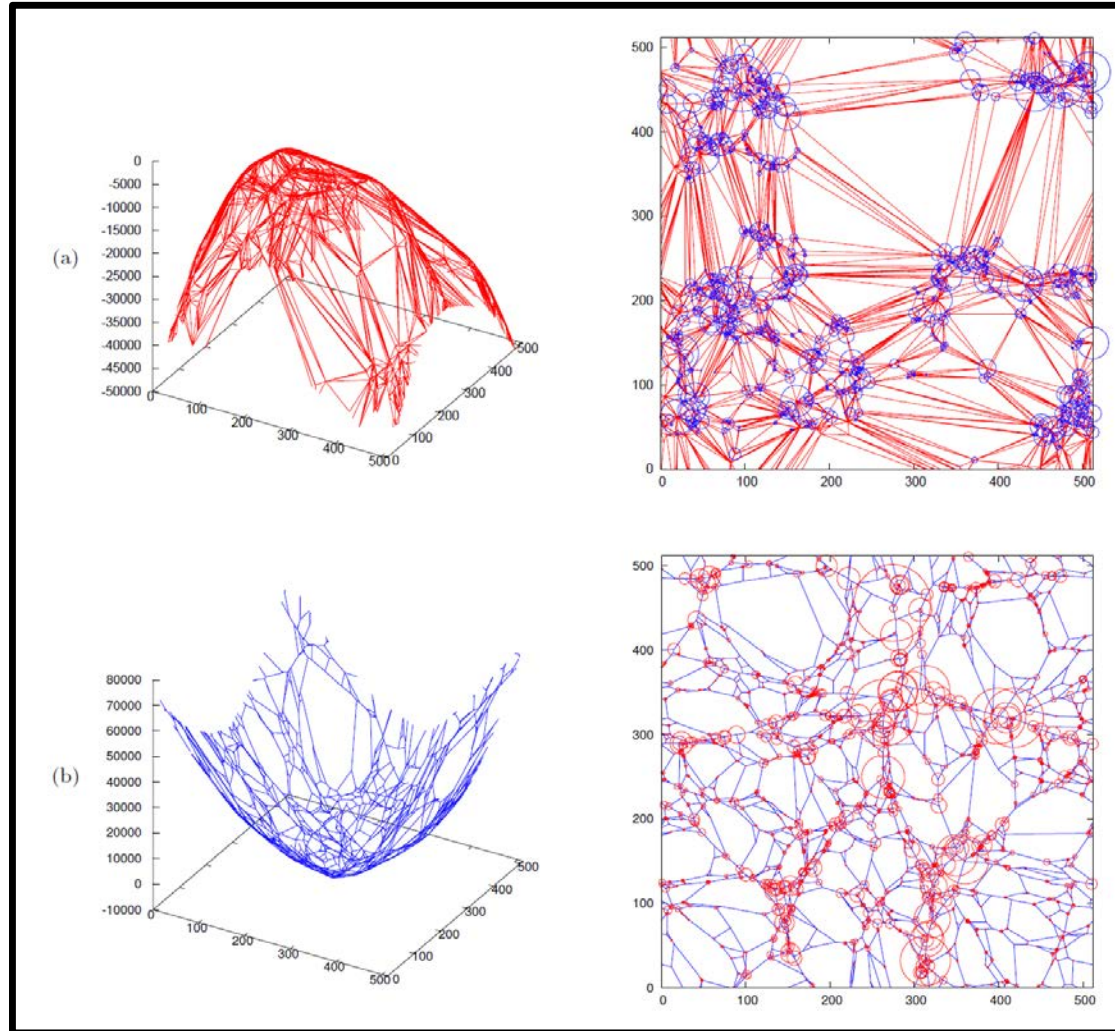


Delaunay tessellation
generated by maxima potential field



Convex Hull

Delaunay-Voronoi: Legendre transform

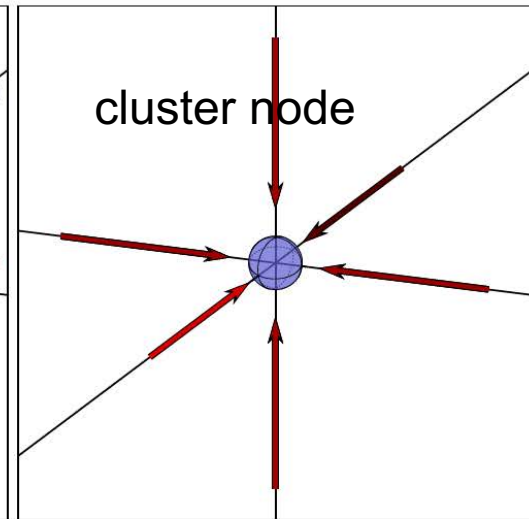
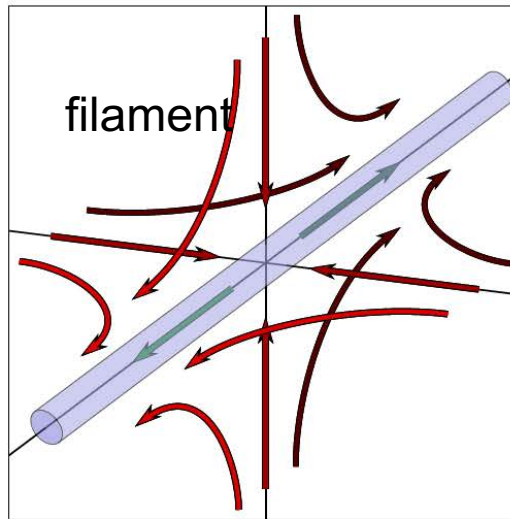
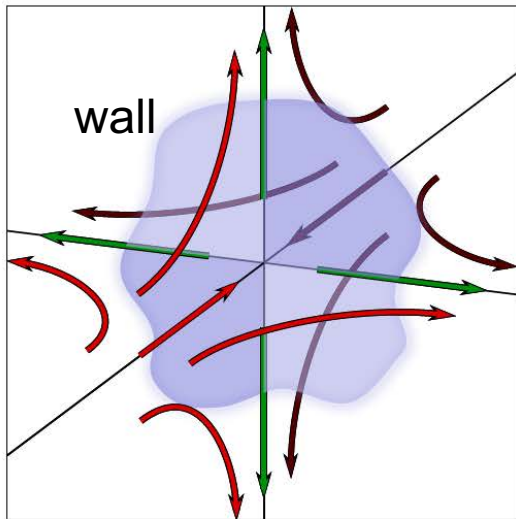


**Delaunay
(weighted)**

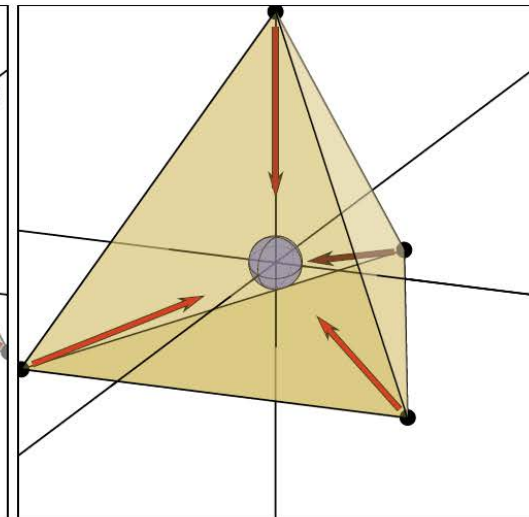
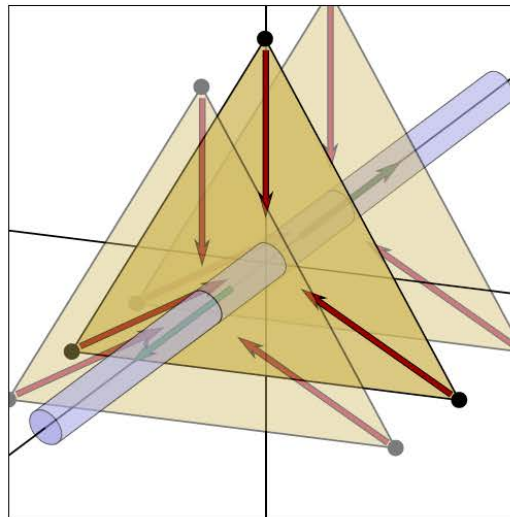
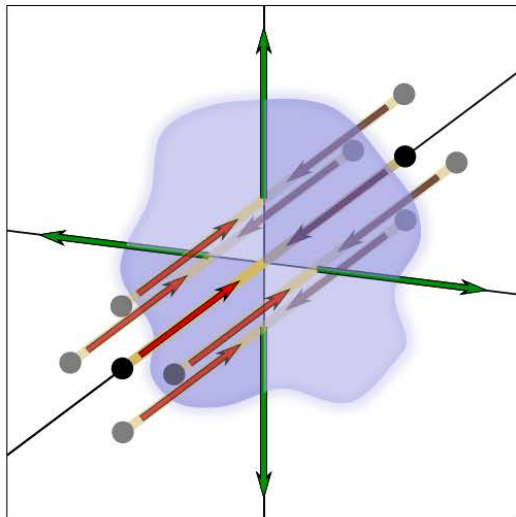
**Voronoi
(weighted)**

$$V_q = \left\{ \vec{x} \in E \left| (\vec{x} - \vec{q})^2 + w_q \leq (\vec{x} - \vec{p})^2 + w_p, \forall \vec{p} \in L \right. \right\}$$

Eulerian vs. Lagrangian weblike geometry



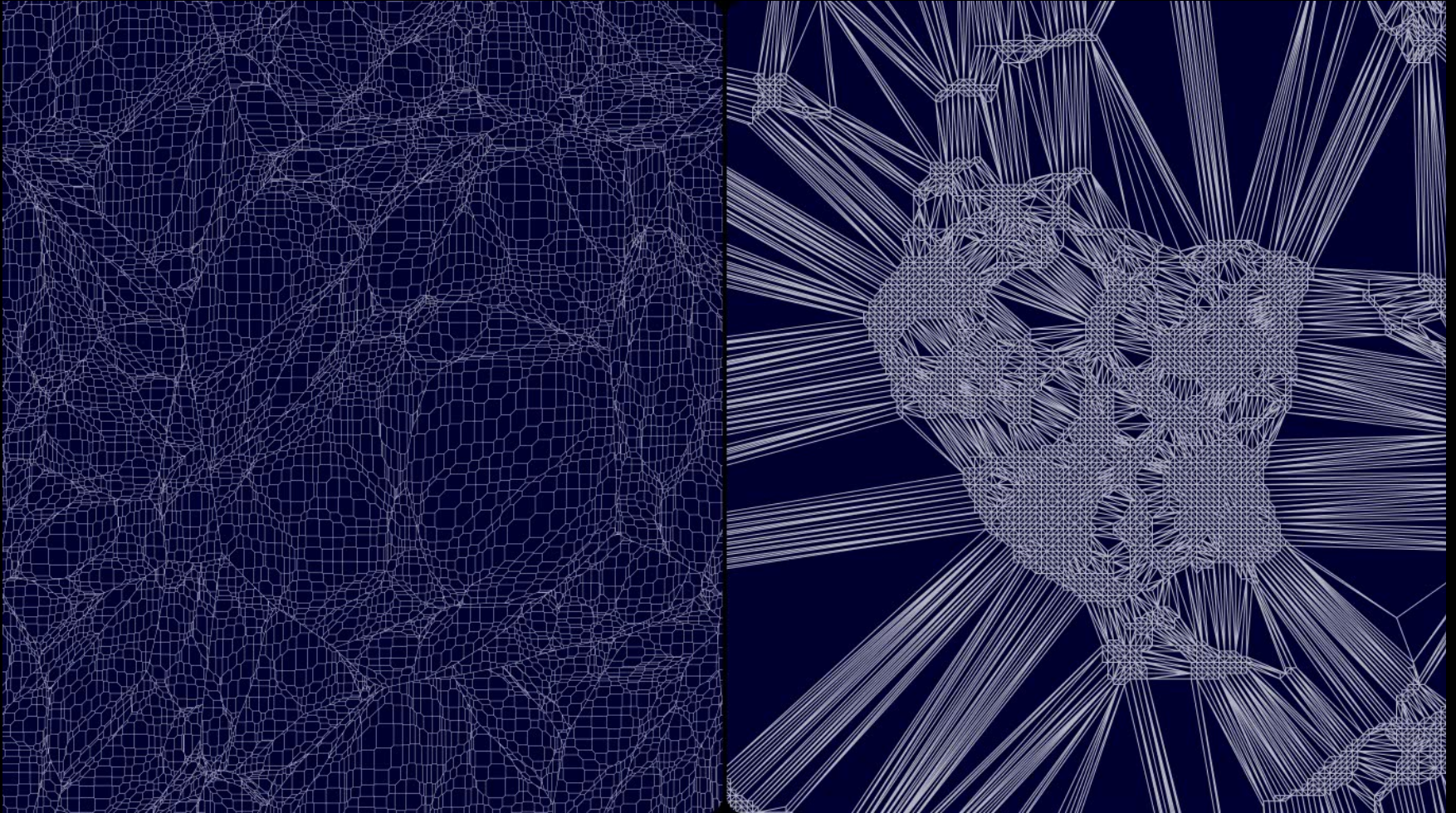
Eulerian



Lagrangian

**Source
regions**

Eulerian – Lagrangian Voronoi - Delaunay

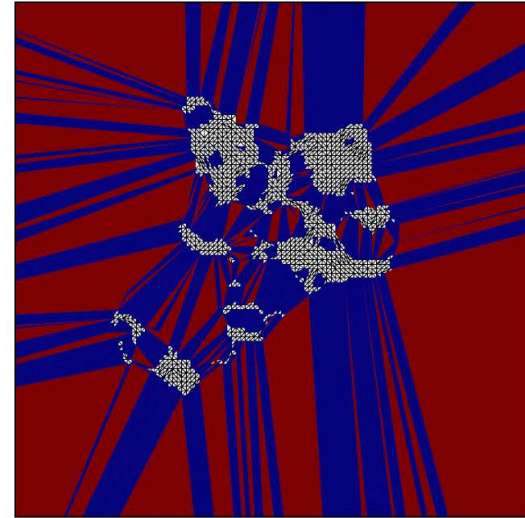
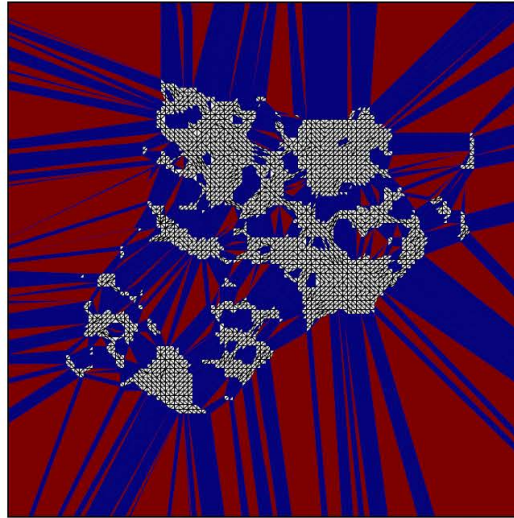
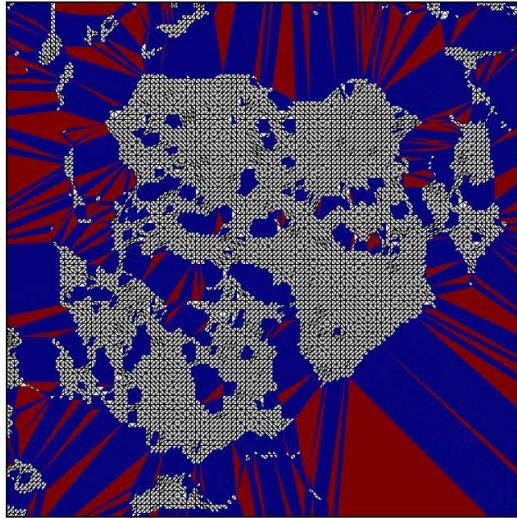


Eulerian – Lagrangian Voronoi - Delaunay



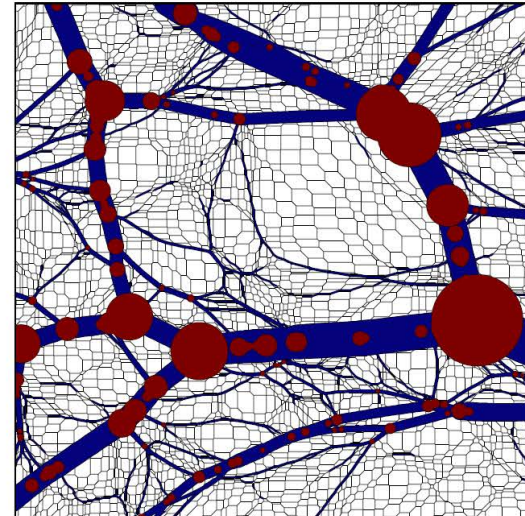
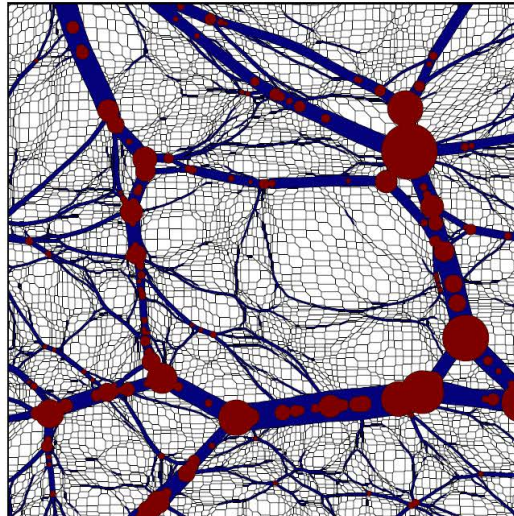
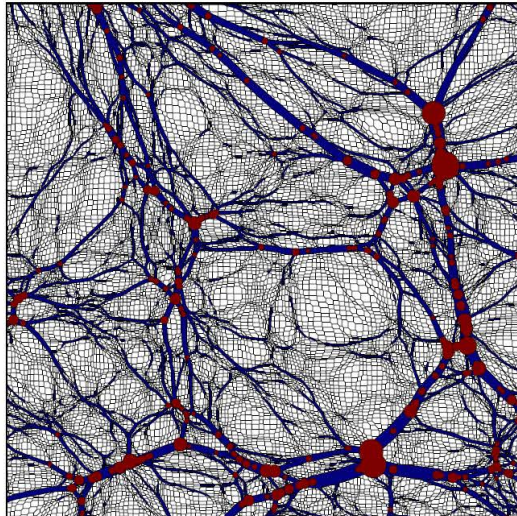
Lagrangian – Eulerian Cosmic Web

Delaunay- Voronoi Tessellations



Lagrangian

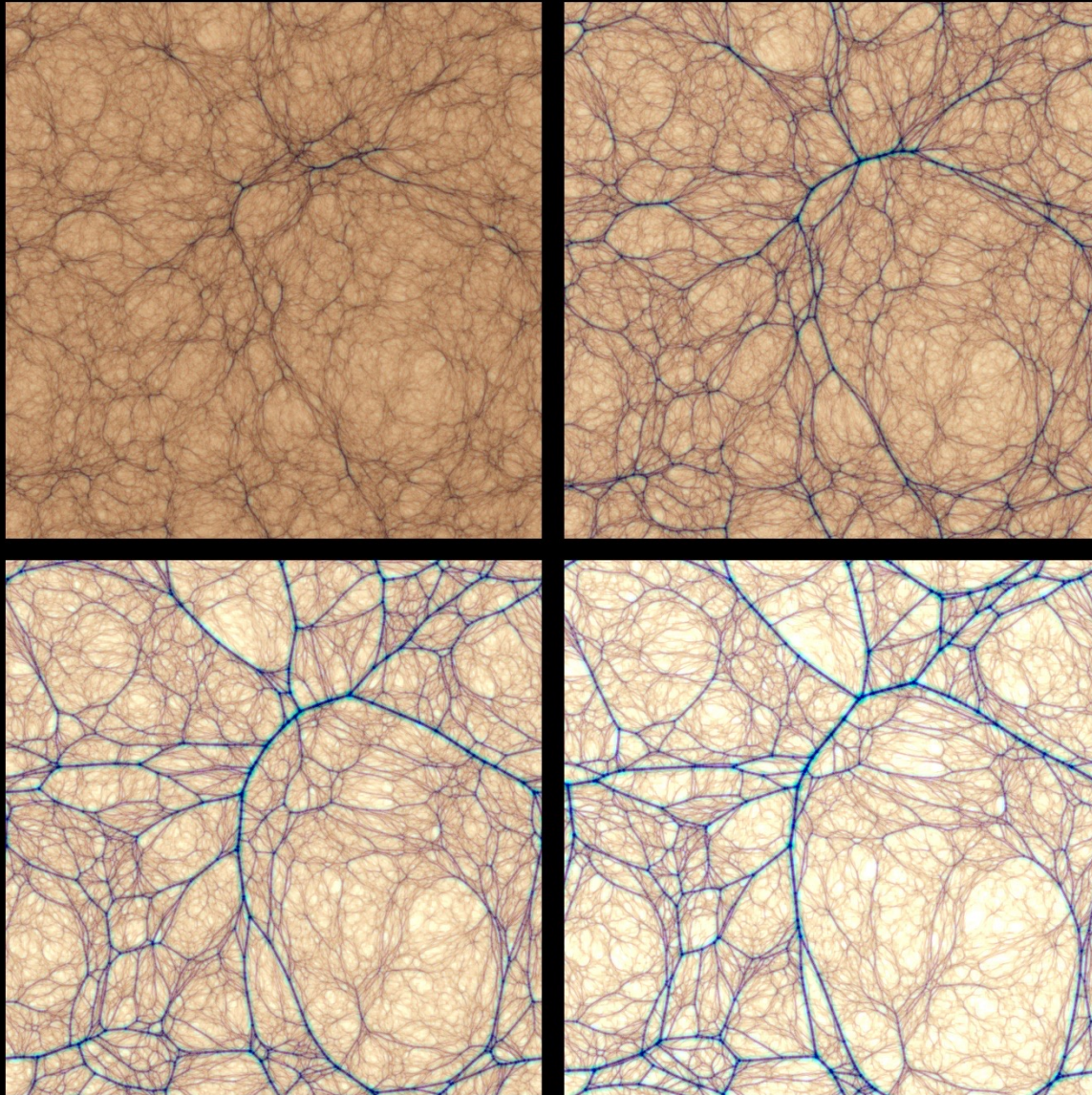
Delaunay
Tessellation



Eulerian

Voronoi
Tessellation

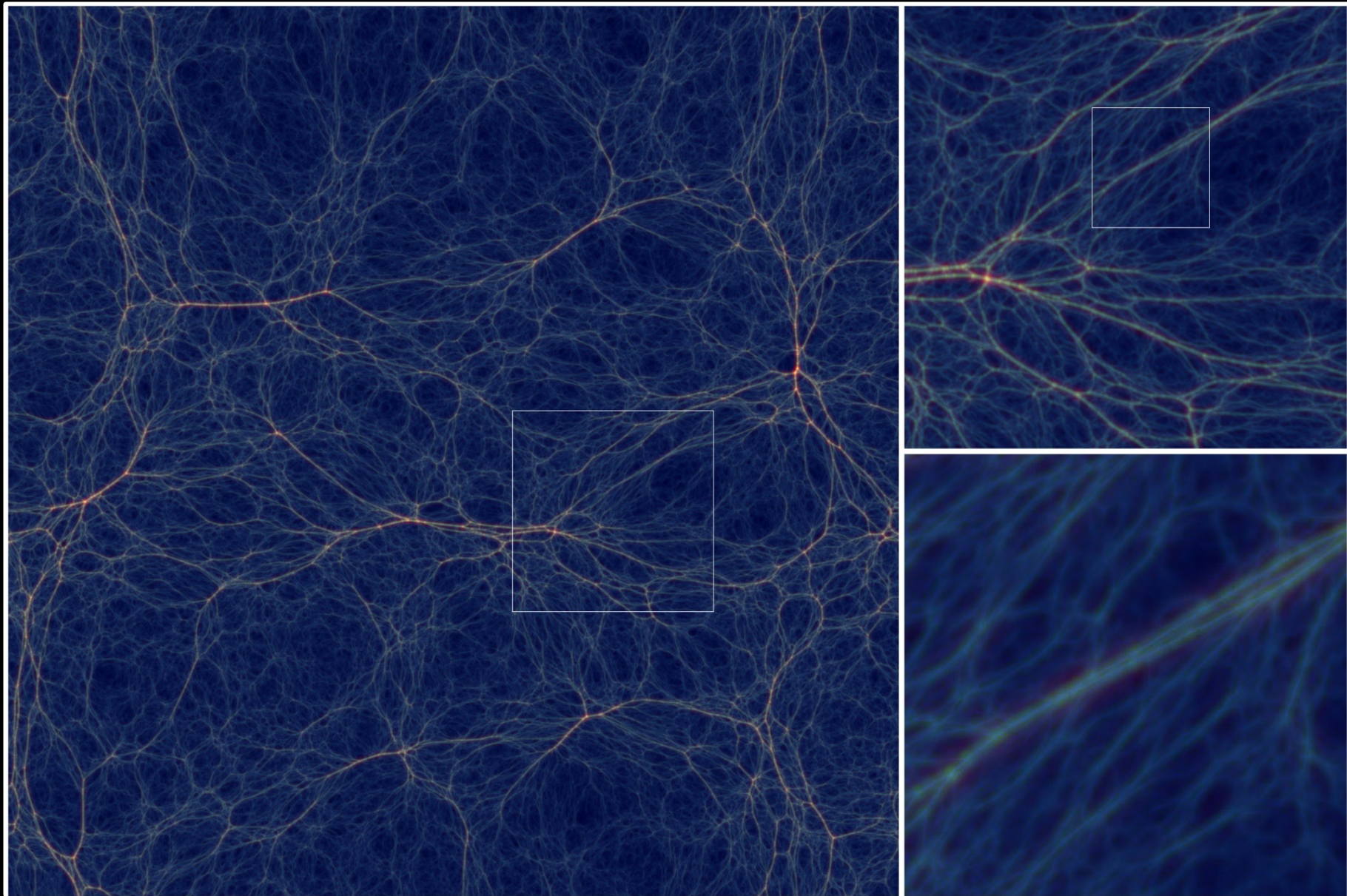
Hierarchical Evolution



The adhesion formalism is ideal for following the hierarchical buildup of the cosmic web:

- Mathematically:
as a result of the evolving parabolic curvature of the (velocity) potential, more features get embedded in singular valleys enclosed between potential and convex hull.
- Physically:
 - Clearly visible is the merging of small filaments into ever larger arteries.
 - at the same time, we see the continuous merging of small voids into larger voids, the evolving soapud of void hierarchy.

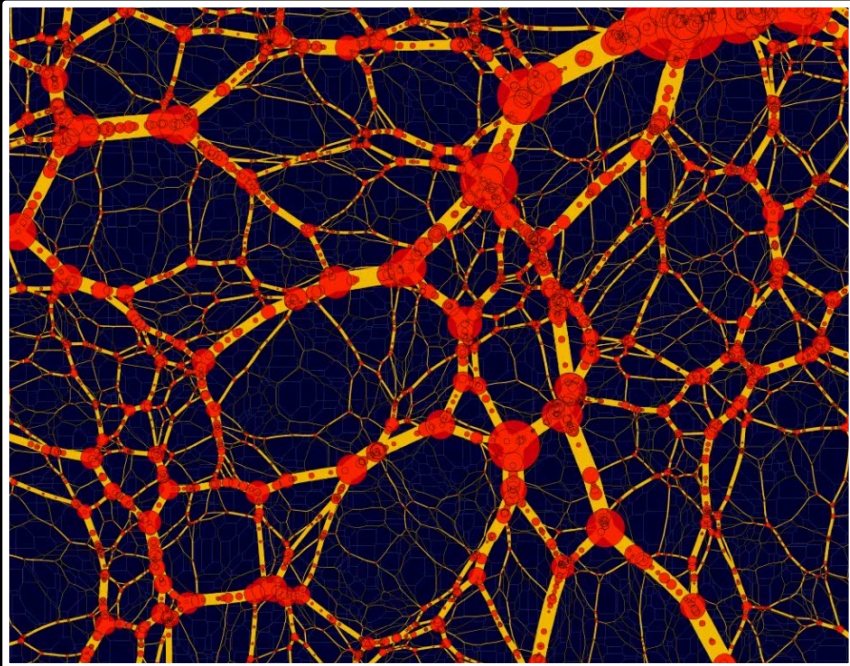
Multiscale Structure



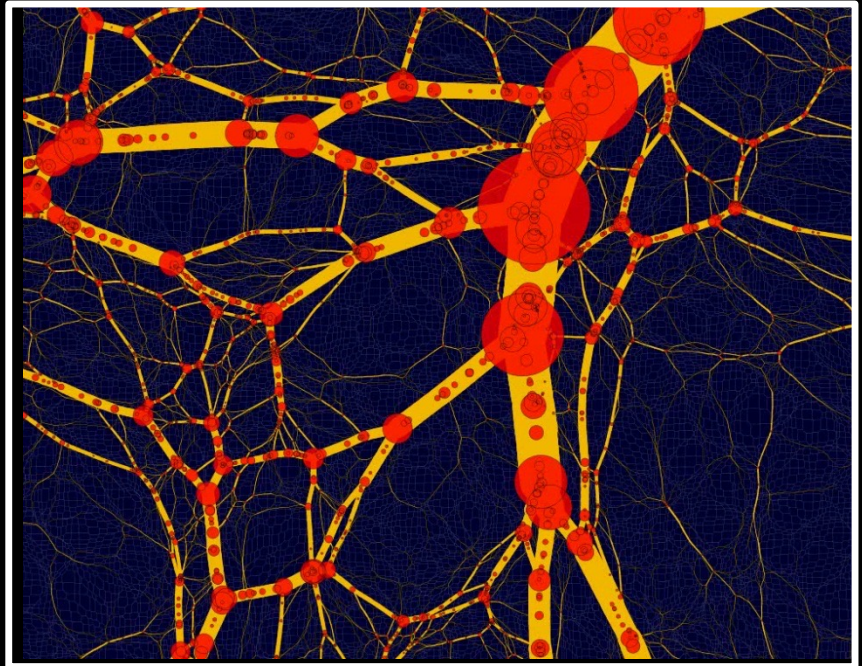
Cosmological Sensitivity Cosmic Web

the morphology of the weblike network is
highly sensitive to the underlying cosmology

$$P(k) \sim k^{-1.5}$$



Hidding 2012/2014



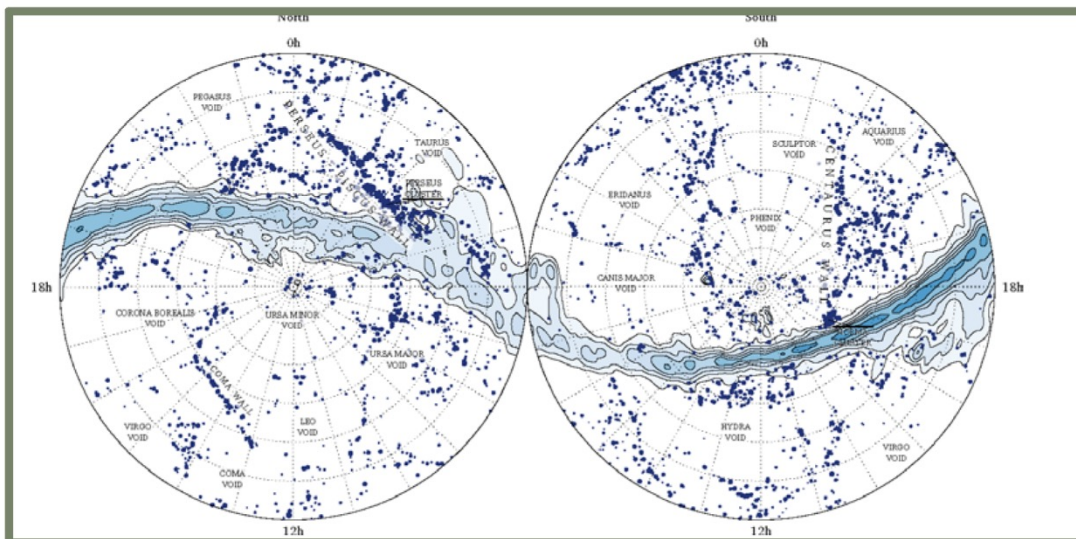
$$P(k) \sim k^{-2.0}$$

A visualization of the cosmic web, showing a complex network of filaments and nodes. The filaments are colored in shades of red, orange, and yellow, while the nodes are darker. The background is black. The entire image is framed by a circular border.

Adhesion Reconstructions :

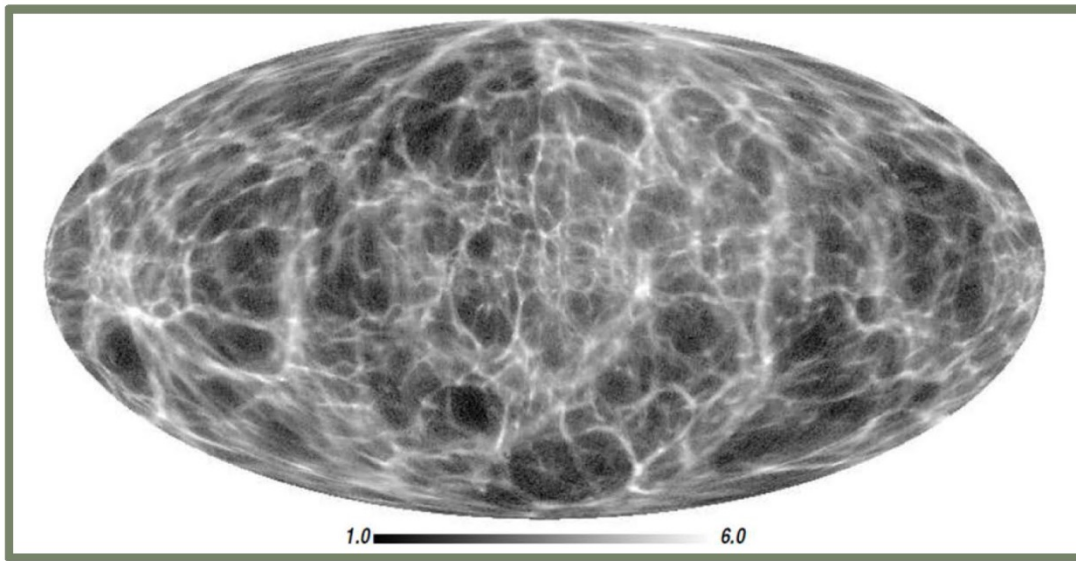
the Cosmic Web in our Local Universe

KIGEN Reconstruction Local Universe ...



2MRS survey sky map
Hidding 2015

Depth: ~50 Mpc



KIGEN (Bayesian)
reconstruction
Local Cosmic Web:

Kitaura

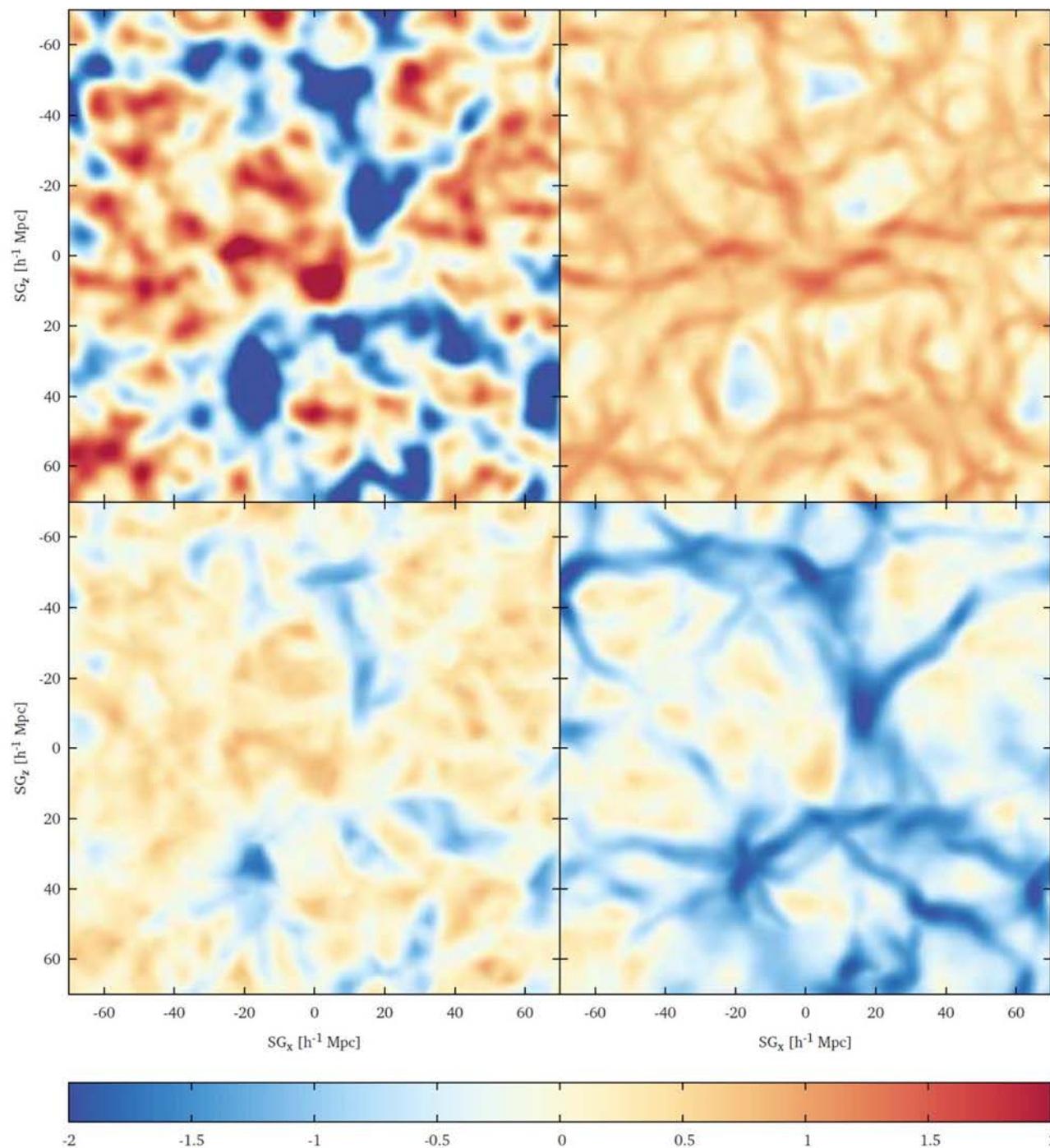
Depth: ~185 Mpc

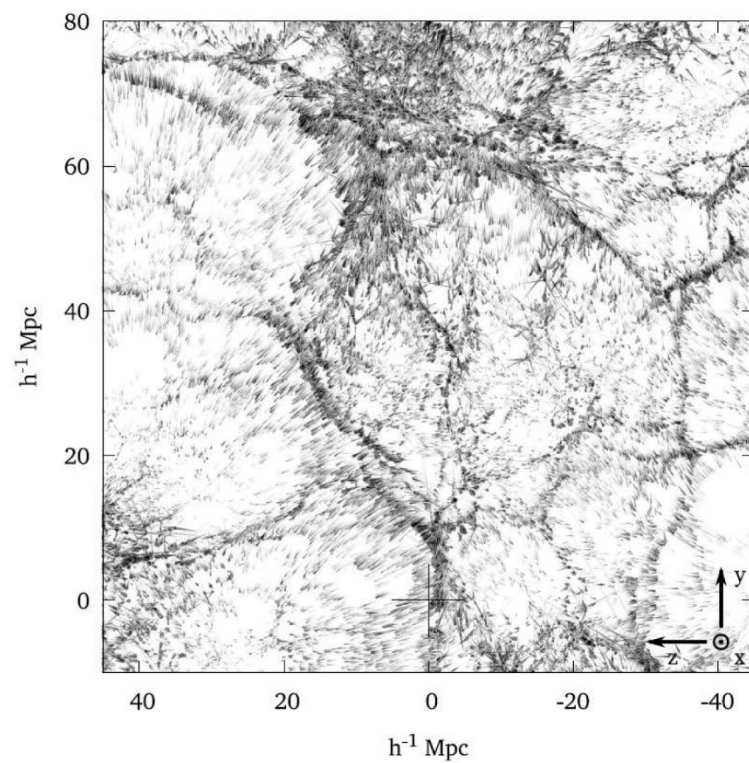
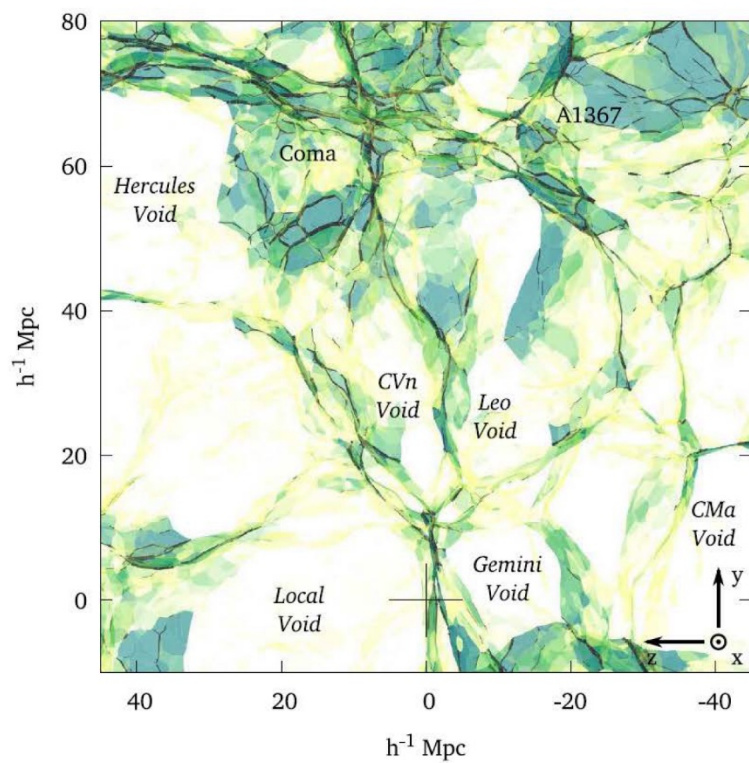
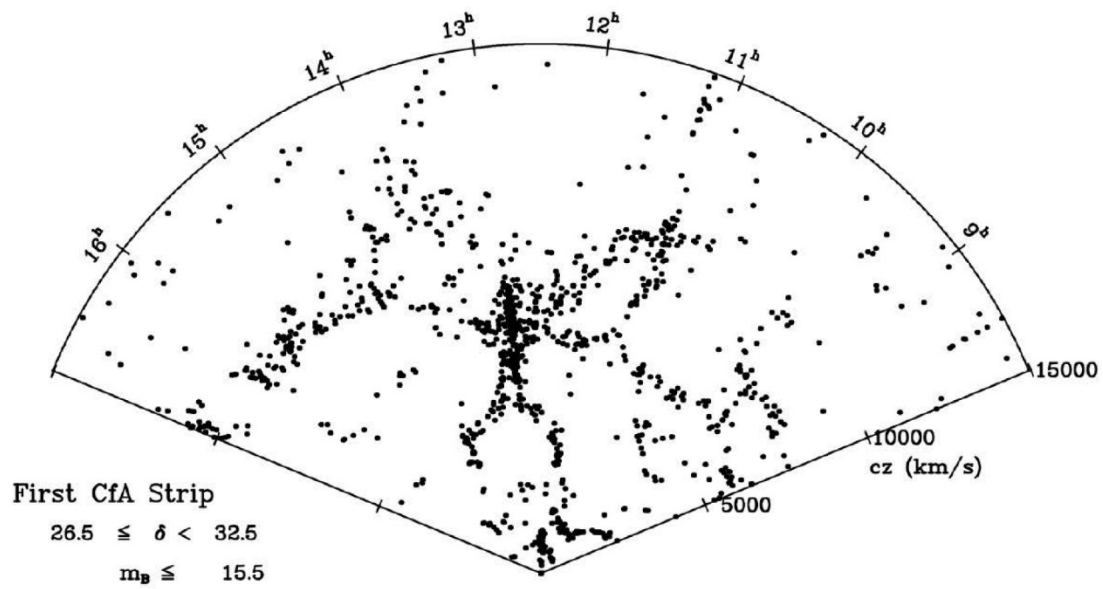
**Initial
Density &
Deformation
Field**

**Local Universe
(SG plane)**

Kitaura & Hess:

**25 KIGEN
constrained
realizations**





Supergalactic Plane

mean adhesion reconstruction

h^{-1} Mpc

60
40
20
0
-20
-40
-60

Coma/A1367

*Leo
Void*

*Coma
Void*

Virgo

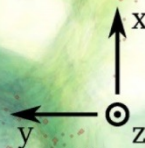
Centaurus

Norma

Perseus-Pisces

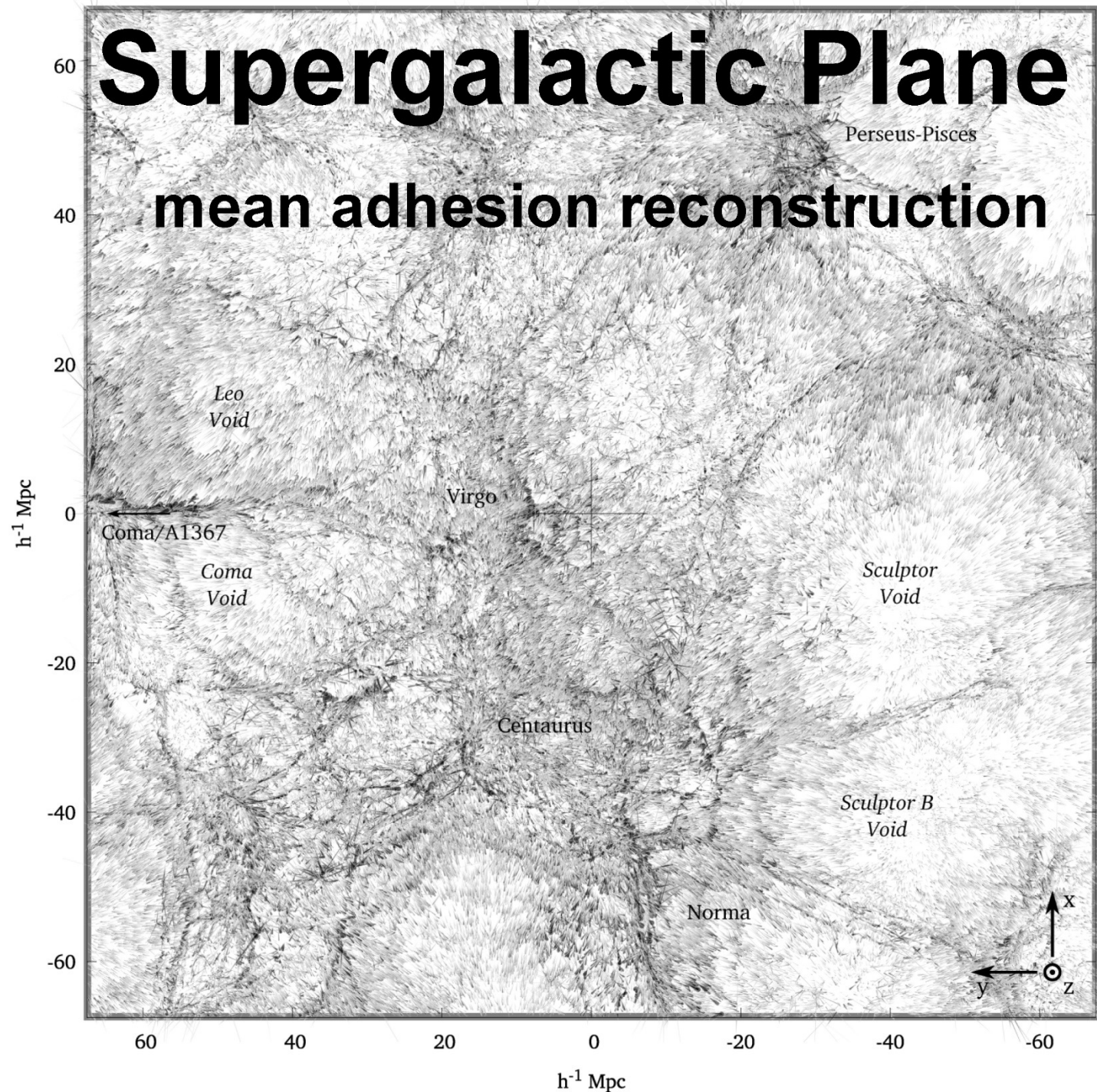
*Sculptor
Void*

*Sculptor B
Void*

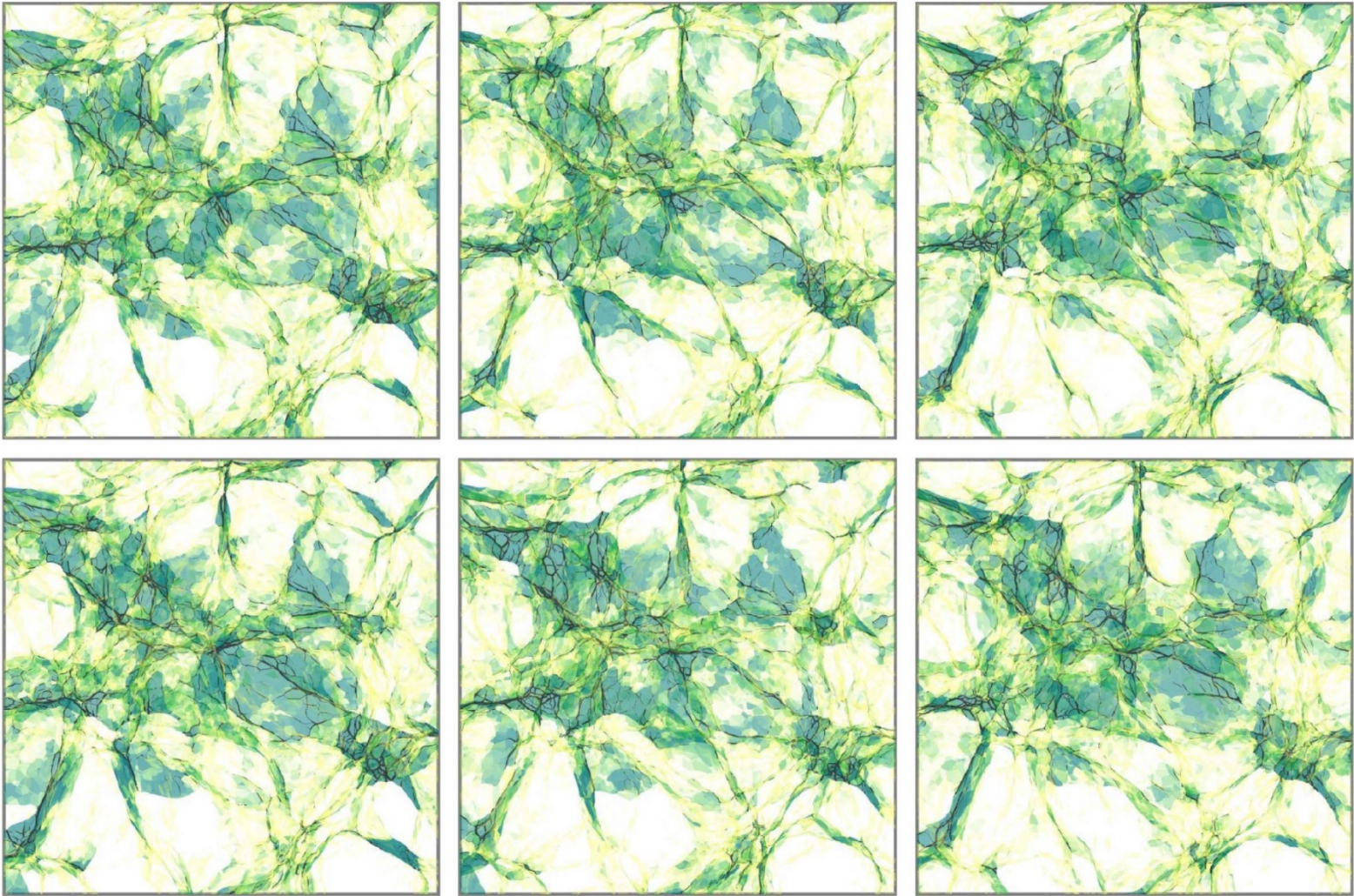


60 40 20 0 -20 -40 -60

h^{-1} Mpc



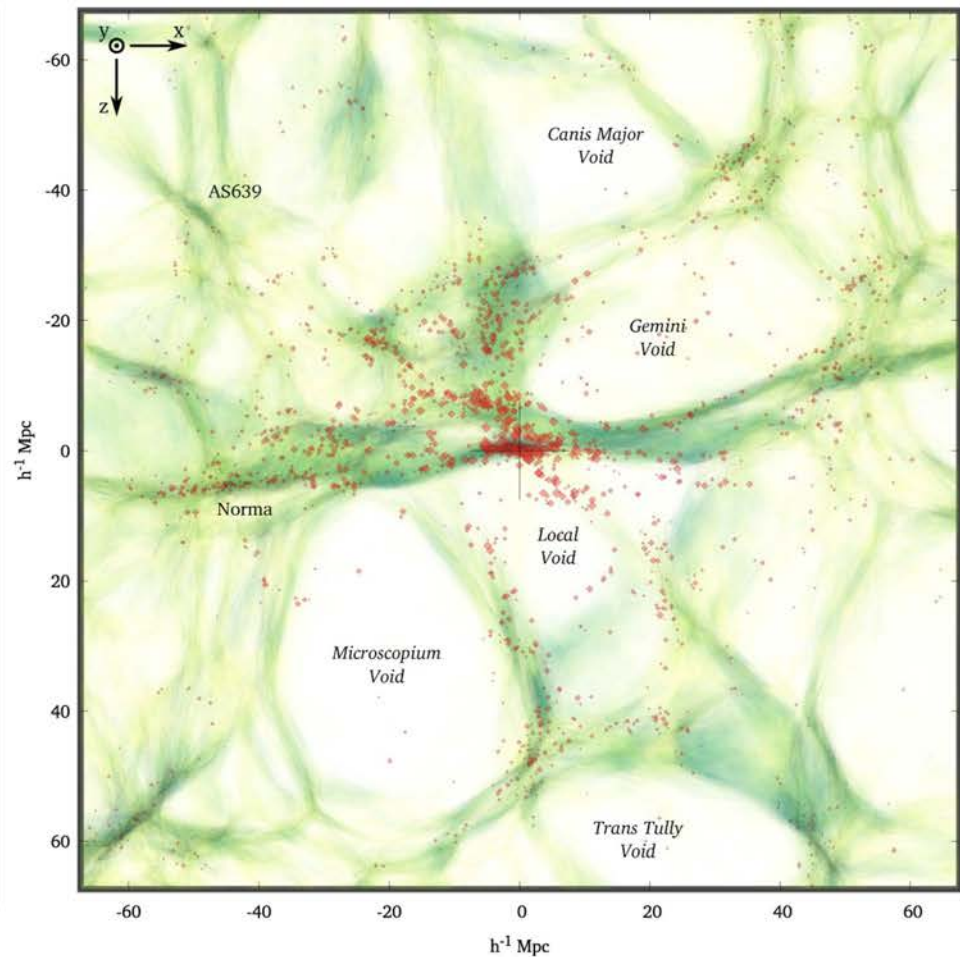
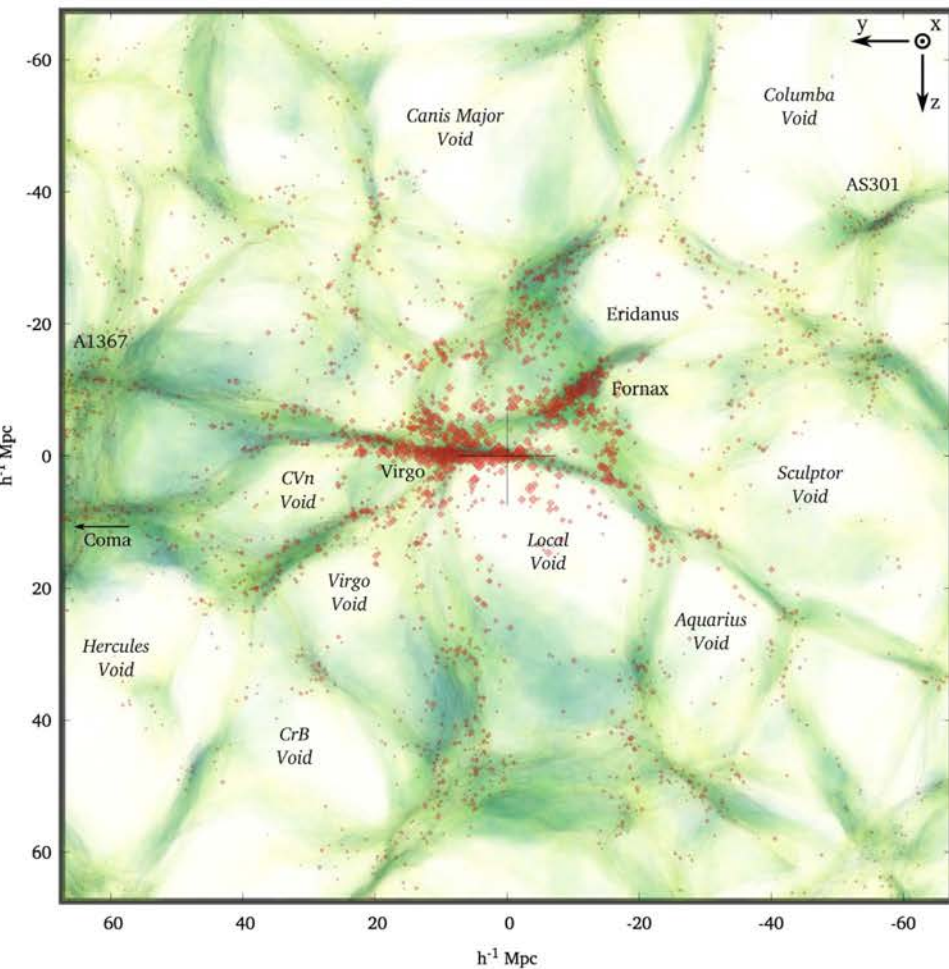
Supergalactic Plane



6 constrained adhesion reconstructions

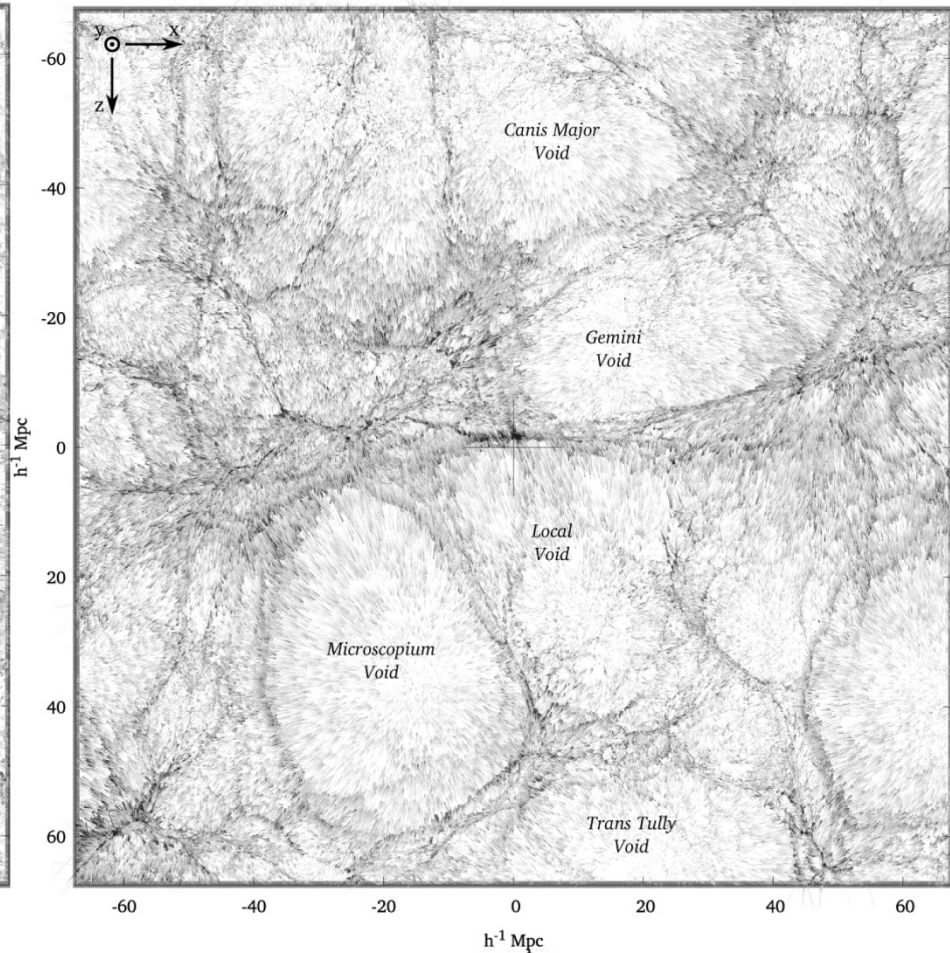
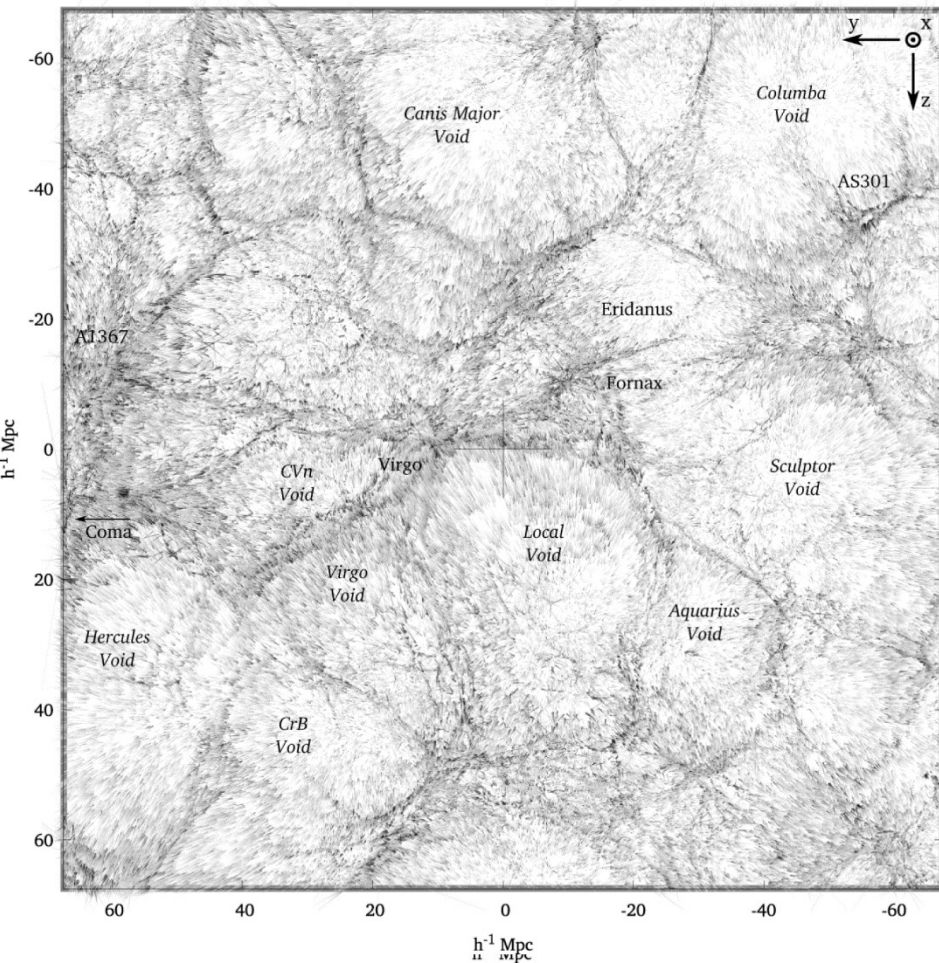
□ Supergalactic Plane

mean adhesion reconstruction



□ Supergalactic Plane

mean adhesion reconstruction



Cosmic Web

&

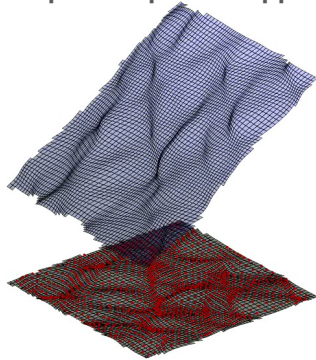
Caustic Skeleton

Cosmic Web – Phase Space Folding

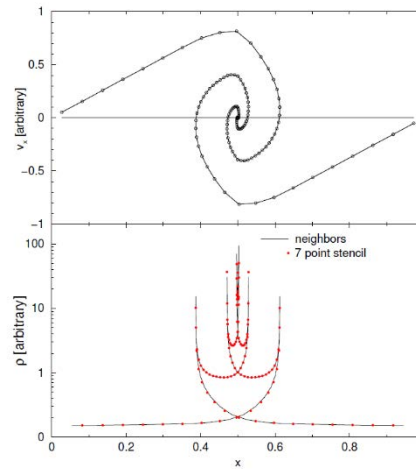


Caustic Skeleton & Phase-Space Wrapping

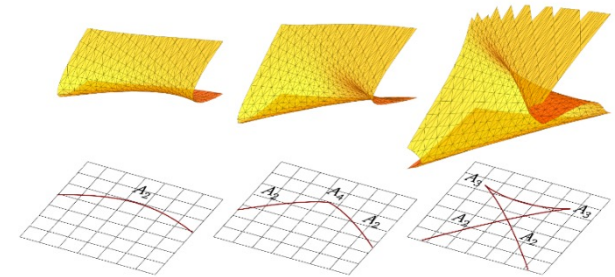
Cosmic Web
phase-space wrapping



Hidding 2014

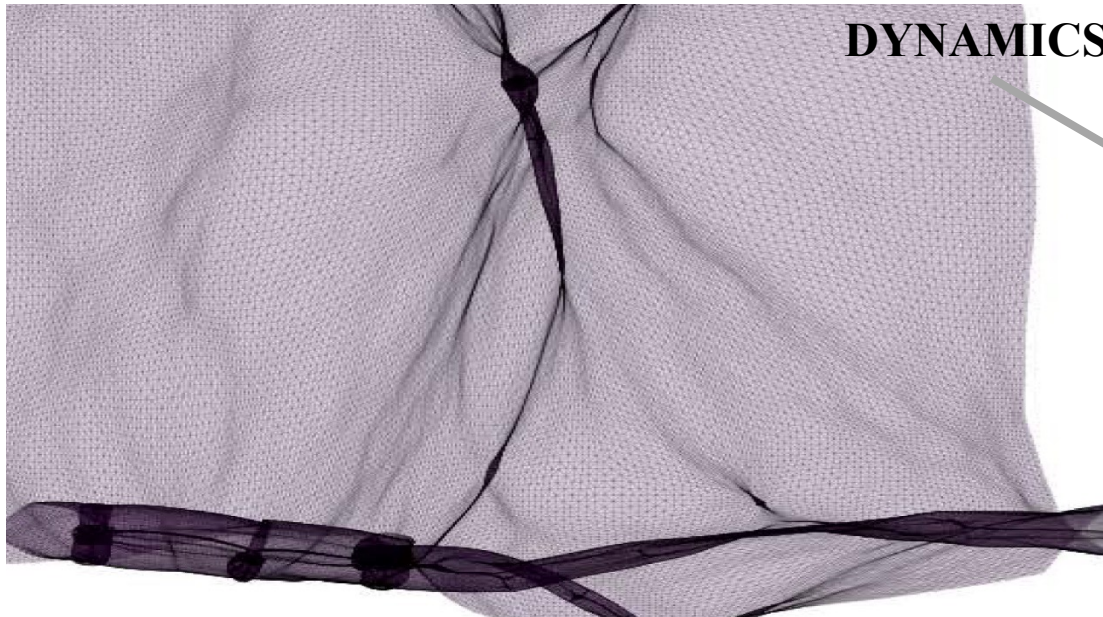


Classification phase-space folding –
Structural Morphology



Feldbrugge, vdW et al., JCAP, 2018
Hidding, Shandarin, vdW, MNRAS, 2014

DYNAMICS



Cosmic Catastrophe Theory:

Lagrangian catastrophe/caustic classification V. Arnold

In Lagrangian space (coordinates q):

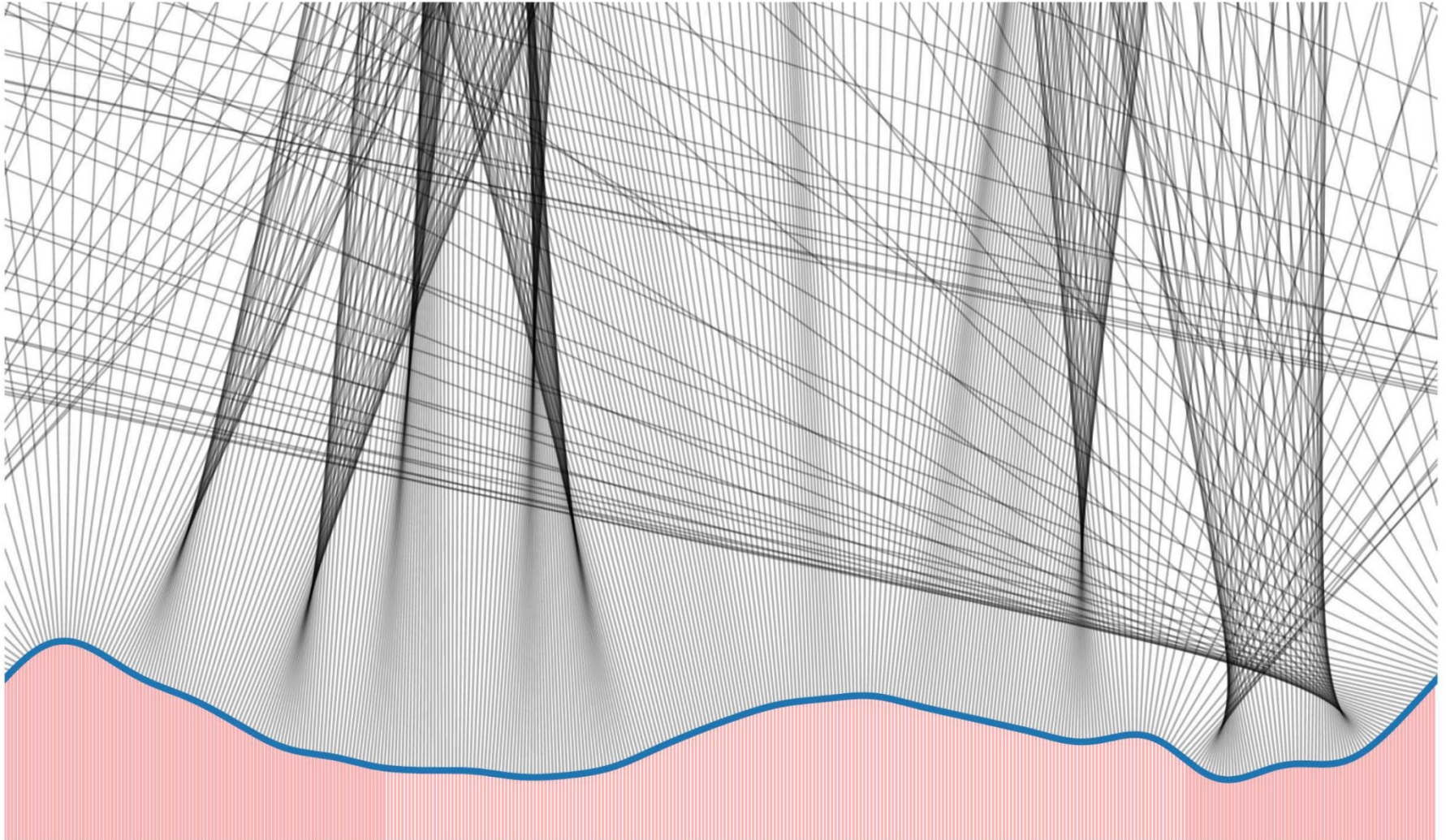
A singularity forms in a manifold M at
location q_s when at q_s ,

- the deformation tensor eigenvalue $\mu_i(q_s)$
- the corresponding eigenvector $\vec{v}_i(q_s)$

when at least one nonzero tangent vector \vec{T}

$$\{1 + \mu_i(q_s)\} \vec{v}_i^*(q_s) \cdot \vec{T} = 0$$

Phase Space Dynamics & Tracks



Deformation, Streaming & Caustics

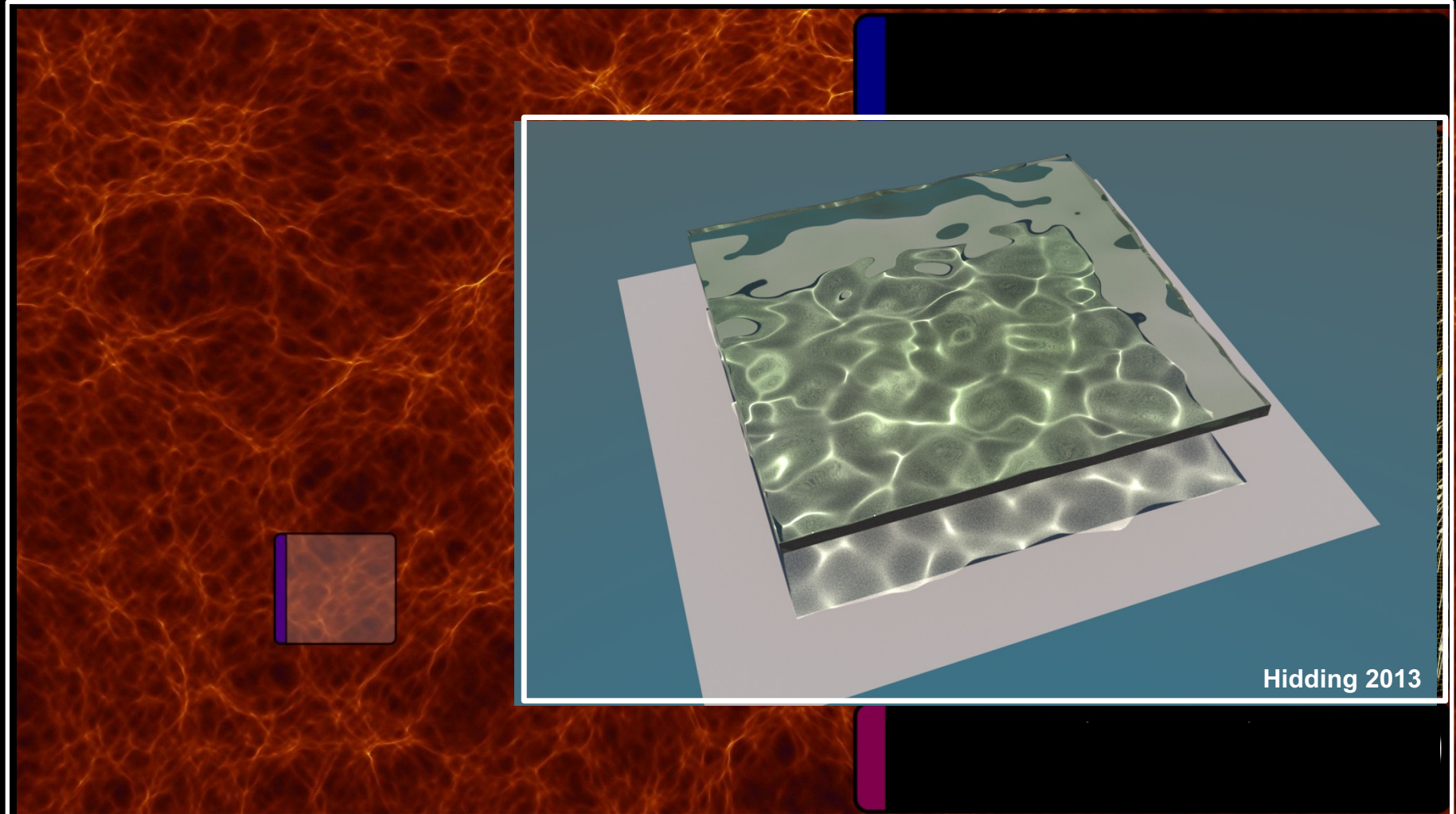
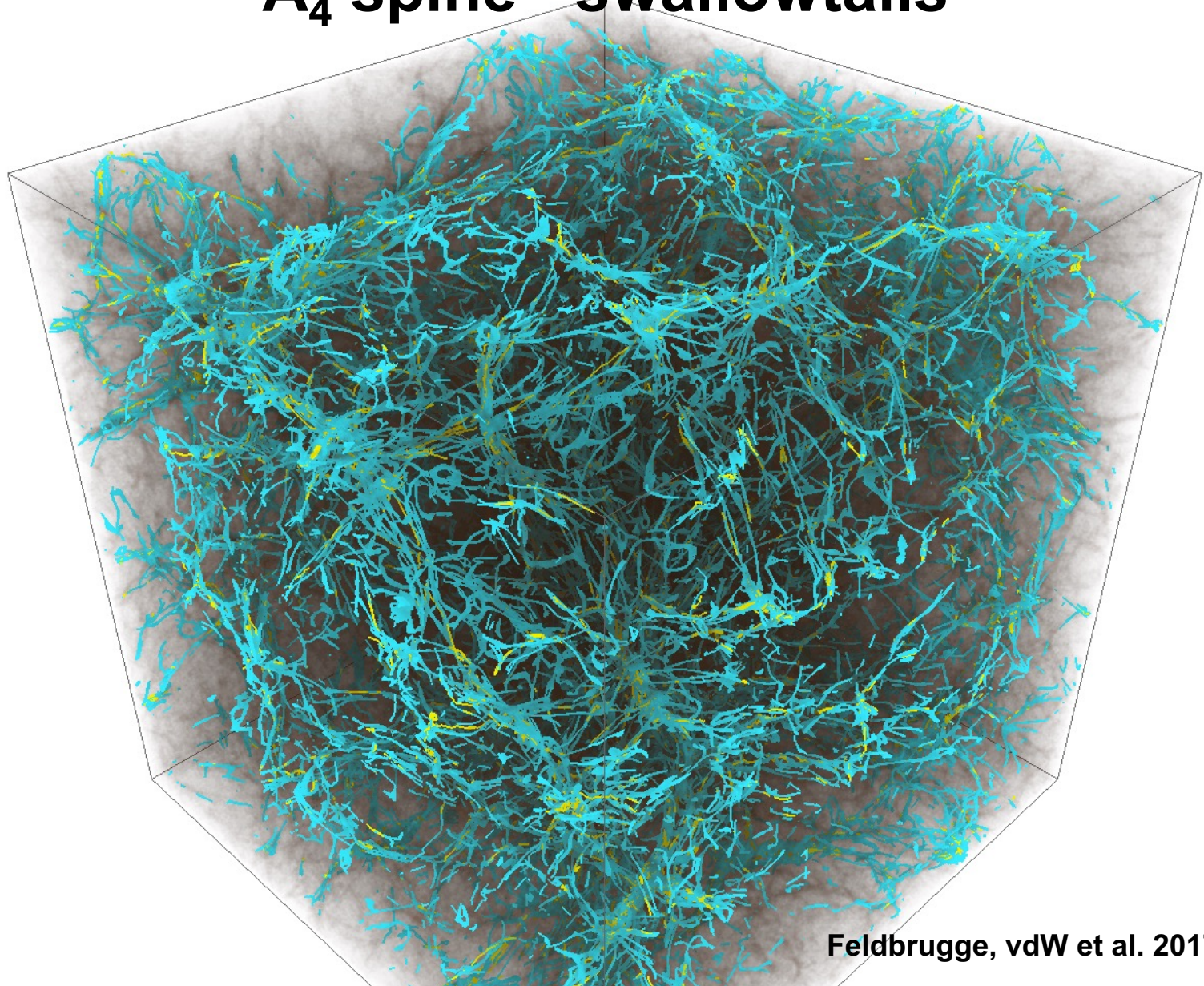


Illustration of the formation of caustics due to
streaming paths of light through deforming medium

Skeleton (3D) Cosmic Web:

A_4 spine - swallowtails



Feldbrugge, vdW et al. 2017b